A Measurement of the ¹⁹Ne Beta Asymmetry & a Determination of $|V_{ud}|$

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Beta Decay Observables



Beta Decay Parameters

Jackson, Treiman and Wyld (Phys. Rev. 106 and Nucl. Phys. 4, 1957)



On-going or planned efforts to measure:

- (1) Decay rates and β -spectra ($G_F V_{ud}, \xi, b$)
- (2) Unpolarized angular correlations $(a_{\beta\nu}, b)$
- (3) Polarized angular correlations $(A_{\beta}, B_{\nu}, b, b_{\nu})$

(4) New program to measure circular polarization asymmetry

Mirrors are isobaric analog mixed decays → two measurements needed to determine both V and A Couplings: Decay Rate + Angular Correlation

Angular Correlations in Nuclei – Polarized Systems

Rather limited set of measurements on polarized nuclei at present-->

Species	Decay	Method	Corr	Corr. unc	Group	
¹⁹ Ne	F/GT	Atomic Beam	Α _β	~2%	Princeton	Complete In 1995
³⁷ K	F/GT	Optical Trap	Α _β	~0.1%	TRINAT-TAMU	ongoing
²¹ Na, ³⁷ K	F/GT	Atomic Beam	σ-Α _β	~0.1%	NSCL	ongoing

any others?

¹⁹Ne (Princeton): in situ polarimetry precision at 1.5%
 ³⁷K (TRINAT-TAMU): in situ polarimetry precision at ~0.1%
 <sup>Motivated to determine mixing ratio...
 Spin-asymmetry (NSCL): running soon, very strong constraints on RHC
</sup>

Many more measurements (on mirrors as well as other systems) planned for unpolarized nuclei..

More experiments coming (see later in talk)!

The β -asymmetry



 $R = R_o(1 + (v/c) P A(E) cosθ)$ β-asymmetry = A(E) in angular distribution of β

$$A_{\beta}(0) = \frac{\rho^2 - 2\rho\sqrt{J(J+1)}}{(1+\rho^2)(J+1)} \qquad \qquad \rho \equiv \frac{C_A M_{GT}}{C_V M_F}$$

Ignoring recoil order terms – just a function of ρ !

Measurement Challenges

 β directional distribution: $1 + P \frac{v}{c} A(E) \cos\theta$ (polarized neutrons)



- $<\cos\theta>$ \longrightarrow Scattering (esp. backscattering)

Spin ratios provide robust 1st order strategy for experiment – "superratio" eliminates detector efficiencies and rate variations

A_{β} in ¹⁹Ne(1/2⁺) \rightarrow ¹⁹F(1/2⁺) Positron Decay

Calaprice group, thesis of Gordon Jones (1995); G. L. Jones, A. Ackerson, M. S. Anderson, F. P. Calaprice, F. Loeser, A. Razaghi, A. R. Young

Hero who finished analysis: D. C. Combs

 $A_{\beta} = -0.0391(14)$ (current)

- $\begin{array}{c} \begin{array}{c} & & & & \\ 1.554 \text{ MeV}, 3/2^{+} & & & \\ & & & \\ 1.554 \text{ MeV}, 3/2^{+} & & \\ & & & \\ & & & \\ 0.110 \text{ MeV}, 1/2 & & \\ & & & \\ 0 \text{ MeV}, 1/2^{+} & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ 0 \text{ MeV}, 1/2^{+} & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & \\ & & & \\ \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \end{array}$ {c} \begin{array}{c} & & & & \\ \end{array} \begin{array}{c} & & & & \\ \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \end{array}{c} \end{array} \begin{array}{c} & & & & \\ & & & & \\ \end{array} \end{array}{c} \end{array}\begin{array}{c} & & & & \\ & & & \\ \end{array} \end{array}{c} \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \end{array}{c} \end{array}{c} \end{array} \begin{array}{c} & & & & \\ & & & \\ \end{array} \end{array}{c} \end{array}{c} \end{array}{c} \end{array}
- Accidental cancellation makes A_β very sensitive to $\rho {:}~\delta A/A \sim 13 d\rho/\rho$

Relaxes demands on systematic error budget! (δA translates into much smaller δρ)

• Critical work sorting out nuclear corrections for mirrors done in 2008 & 2009:

Severijns et al., PRC **78**, 055501 (2008) Naviliat and Severijns, PRL **102**, 142302 (2009) T_{1/2} to ground state: 17.2818(94)

2.727 MeV, 1/2

K.E. max = 2.216 keV

$$\begin{split} \mathsf{M}_{\mathsf{F}} &= 1\\ \mathsf{f}_{\mathsf{A}}/\mathsf{f}_{\mathsf{V}} &= 1.0143(29)\\ (1+\Delta_{\mathsf{R}}) &= 1.02361(38)\\ (1+\delta_{\mathsf{R}}) &= 1.01533(12)\\ (1+\delta_{\mathsf{NS}}) &= .9948(4) \end{split}$$

Princeton/Berkeley Polarized Atomic Beam Apparatus

(State of the Art until well after 2000)



~2000 – 3000 polarized decays/sec in cell

Asymmetry



P=+1.6015(29) for convention of Severijns et al

Error Budget ¹⁹Ne

Systematic	Correction (10 ⁻⁴)	Uncertainty (10 ⁻⁴)
Monte Carlo Corrections:		
Above threshold in both detectors:		
Backscatter correction	+14.5	±3.6
Energy loss correction	-2.0	± 0.5
Above threshold in a single detector:		
Backscatter correction	+3.0	± 0.8
Energy loss correction	-0.9	±0.2
Below threshold in both detectors:	-0.5	±0.1
Polarization	_	+5.7 -0.0
Spin relaxation	+5.3	±5 .3
Energy non-linearity	-	± 0.5
Dead time	-0.5	± 0.4
Pileup	-0.6	± 0.4
Background subtraction	+0.2	±0.2
Statistical	_	± 3.0
Total	+18.5	+9.2 -7.2

$\delta A/A = 2.47\%$

(previous value, 3.9%)

Not limited by statistics

Error Budge	Systematic errors:		
Systematic	Correction (10 ⁻⁴)	Uncertainty (10 ⁻⁴)	the usual suspects.
Monte Carlo Corrections:			
Above threshold in both detectors:			
Backscatter correction	+14.5	±3.6	
Energy loss correction	-2.0	± 0.5	
Above threshold in a single detector:			Scattering corrections (2)
Backscatter correction	+3.0	± 0.8	
Energy loss correction	-0.9	±0.2	
Below threshold in both detectors:	-0.5	± 0.1	
Polarization	_	+5.7 -0.0	
Spin relaxation	+5.3	±5 .3	Polarization (1)
Energy non-linearity	-	±0.5 ┥	——————————————————————————————————————
Dead time	-0.5	± 0.4	
Pileup	-0.6	± 0.4	
Background subtraction	+0.2	±0.2 🗲	——Background Subtraction
Statistical	_	± 3.0	
Total	+18.5	+9.2 -7.2	_

 $\delta A/A = 2.47\%$

Error Budget ¹⁹Ne



 $\delta A/A = 2.47\%$

Error Budget ¹⁹Ne



 $\delta A/A = 2.47\%$

Dustin Combs thesis: re-analysis of scattering corrections, including backscattering reconstruction

Scattering Correction

Strategy: use timing to reconstruct backscatters which hit one detector (e.g. D1) and then scatter into the second (D2) – use T1-T2 to determine initial direction of beta!



Full PENELOPE model of both beta-asymmetry timing spectrum and timing calibration measurements, together with detector model including charge transport of quasiparticles in Si Overlap region results in Errors in assignment of dir!

 $\Delta A_{\beta}/A_{\beta} = 3.8(0.9)\%$

PENELOPE v2002 – vetted with direct tests and in the UCNA experiment Backscattering most challenging – 25% uncertainty assigned to MC results

Error Budget ¹⁹Ne



Gordon Jones did an excellent job of $\delta A/A = 2.47\%$ optimizing the performance of the polarizer, a device in use for almost 40 years – expected polarization was > 99%

Polarization (old school)



Set conservative lower limit on polarization by assuming background completely depolarized

Error Budget ¹⁹Ne



$$\delta A/A = 2.47\%$$





Status of ¹⁹Ne

- Overall uncertainty $dA_{\beta}/A_{\beta} = 2.47\%$
- Leading uncertainty from polarization (1.5% from beam polarization, 1.1% from depolarization), next is scattering (0.9%)
- Lifetime uncertainty ~0.02%
- Results in $\delta V_{ud}/V_{ud} = 0.16\%$ for ¹⁹Ne alone (superior to the PDG 2018 neutron value)
- Uncertainty from A now comparable to theory uncertainties $(f_A/f_v)!$

Theory Needs

$$f_V t (1 + \delta'_R) \left(1 + \delta^V_{NS} - \delta^V_C \right) \\= \frac{K}{G_F^2 V_{ud}^2} \frac{1}{\left| M_F^0 \right|^2 C_V^2 \left(1 + \Delta^V_R \right) \left(1 + \frac{f_A}{f_V} \rho^2 \right)},$$

One other quantity that depends weakly on a shell-model calculation is the ratio f_A/f_v (column 4 in Table VIII). Here a modest shell-model calculation is sufficient. We can also use these shell-model calculations to determine the relative sign of the Fermi and Gamow-Teller matrix elements, which can then be taken as the sign of ρ in Eq. (22). Finally, the resulting Ft mirror values and corresponding values for ρ (using Ft^{0+ -0+} = (3071.4 ± 0.8) s [25], and assigning an error of 20% to the calculated deviations of f_A/f_v from unity) are recorded in Table IX.



 ρ^2 a factor of 4 or more greater than other mirrors except neutron (where f_{a}/f_{v} correction is order 10⁻⁵)

Next Generation Angular Correlations

How has the field moved forward to improve?

- Ion trap measurements of the beta-neutrino correlations, $a_{\beta\nu}$
- Laser trap measurements of A_{β} , $a_{\beta\nu}$

How to improve precision:

- Produce highly localized, "massless" source (no cell) lon and optical traps ideal
- Reduce/eliminate scattering effects from grids, apertures, detectors
 Use position sensitive reconstruction, low mass,

low Z components

• Eliminate backgrounds

Two-stage trapping, pure samples, coincidence signals, event reconstruction

Common elements of current expts

When are we projected to be ready for an significant jump in the precision of these measurements?

Now!

Example:

Laser-trapped species include Alkali metals (³⁷K) and metastable noble gas atoms (¹⁹Ne)





Over order of magnitude improvement relative to Princeton Measurement

In situ polarimetry to 0.05% (!)

Over order of magnitude improvement relative to Princeton Measurement

Source		Correction	Uncertainty
Systematics Background β scattering®		1.0014 1.0230	0.0008
Trap(σ+ vsσ) { position (typ ≲ sail velocity(ty temperature (ty	$(\pm 20 \ \mu m)$ $p \lesssim \pm 30 \ \mu m/ms)$ $p \lesssim \pm 0.2 \ mK)$	0.0004 0.0005 0.0001
Si-strip { radi ener thre	us*(15.5 ^{+3.5} mm) gy agreement (±30 shold (60 → 40 keV	σ → ±5σ) 7)	0.0004 0.0002 0.0001
Shakeoff elec	tron TOF region (±	$\pm 3.8 \rightarrow \pm 4.6$ ns)	0.0003
Thicknes ses {	SiC mirror ^a ($\pm 6 \mu$ m) Be window ^a (± 23 Si-strip ^a ($\pm 5 \mu$ m)	m) μm)	0.0001 0.000 09 0.000 01
Scintillator or Scintillator th Scintillator ca	ly vs $E + \Delta E^{*}$ reshold (400 \rightarrow 100 libration (±0.4 ch/	00 keV) 'keV)	0.0001 0.000 03 0.000 01
Total systemati Statistics Polarization	25	1.0088	0.0013 0.0013 0.0005
Total		1.0338	0.0019

^aDenotes sources that are related to β^+ scattering.

Leading systematic corrections come from scattering and backgrounds

Total precision improved by an order of magnitude

Technology exists to push ¹⁹Ne to precision levels competitive with superallowed decays!

Implications

- Incredible progress made on scattering corrections and polarimetry open the door to sub-.1% measurements on ¹⁹Ne (being pursued by Ron's group at HUJI), ³⁷K and ²¹Na! This will certainly impact the global beta decay landscape...
- Theory input is also needed. In the short run, the precision of f_A/f_V must be specified over an order of magnitude more precisely for ¹⁹Ne. In the longer run, a deeper understanding of the nuclear structure corrections are needed to convincingly establish precision levels at the 0.02% level and below!

Would high precision beta spectra help constrain NS models?