Standard Model β spectra

Leendert Hayen ACFI Workshop, November 1st 2018

IKS, KU Leuven, Belgium

Motivation

Beta spectrum shape

Current status

Challenges

Motivation

Three basic questions:

What's our goal?

Understand Standard Model & Go Beyond

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Where to look for it? Quirky weak interaction! Three basic questions:

What's our goal? Understand Standard Model & Go Beyond

Where to look for it? Quirky weak interaction!

How to do this?

- β Spectrum shape
- 1. Direct BSM sensitivity
- 2. Enters into reactor anomaly

General Hamiltonian

$$\mathcal{H} = \sum_{j=V,A,S,P,T} \langle f | \mathcal{O}_j | i \rangle \langle e | \mathcal{O}_j [C_j + C_j \gamma_5] | \nu \rangle + h.c.$$

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In Standard Model only $V-A \rightarrow$ where are the **others**?

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QCD influences \rightarrow induced currents, influenced through <code>nuclear structure?</code>

Typical BSM searches through correlations

$$\frac{d\Gamma}{dE_e d\Omega_e d\Omega_\nu} \propto 1 + a_{\beta\nu} \frac{\vec{p_e} \cdot \vec{p_\nu}}{E_e E_\nu} + b_F \frac{m_e}{E_e} + A \frac{\vec{p_e}}{E_e} \langle \vec{I} \rangle + \dots$$

Measure effective correlations

$$ilde{X} = rac{X}{1 + b_F \langle rac{m_e}{E_e}
angle}$$



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Sensitivity comes from b_F

$$b_{F} = \pm \frac{1}{1+\rho^{2}} \left[\operatorname{Re}\left(\frac{C_{S} + C_{S}'}{C_{V}}\right) + \rho^{2} \operatorname{Re}\left(\frac{C_{T} + C_{T}'}{C_{A}}\right) \right]$$

because it's linear in coupling constants

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because it's linear in coupling constants \rightarrow measure β spectrum directly & fit for $1/E_e$ Exploring the Standard Model and Beyond via the allowed β spectrum shape:

$$rac{dN}{dE_e} \propto 1 + rac{b_{ extsf{Fierz}}}{E_e} + b_{WM}E_e$$

*b*_{Fierz}: Proportional to scalar (Fermi) and tensor (Gamow-Teller) couplings

 b_{WM} : Weak Magnetism (main induced current), poorly known for A > 60, forbidden decays

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This requires knowledge of the theoretical spectrum shape to $\leq 10^{-3}$ level!

Beta spectrum shape

Beta Spectrum Shape

3-body decay

$$P(E_e) = (E_0 - E_e)^2 E_e p_e \approx (E_0 - E_e)^2 E_e^2$$

Beta Spectrum Shape

3-body decay

$$P(E_e) = (E_0 - E_e)^2 E_e p_e pprox (E_0 - E_e)^2 E_e^2$$



Ellis & Chadwick, 1914

Weak Hamiltonian is modified

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Large scale gap to cross

 $\mathsf{Quark} \to \mathsf{Nucleon} \to \mathsf{Nucleus} \to \mathsf{Atom} \to \mathsf{Molecule}$

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Whole slew of approximations introduced

Starting from the Standard Model $SU(2)_L imes U(1)_Y$ EW sector

$$\mathcal{M} = rac{g^2}{8} V_{ud} ar{u} \gamma^\mu (1 - \gamma^5) d rac{g_{\mu
u} - q_\mu q_
u / M_W^2}{q^2 - M_W^2} ar{e} \gamma^
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Since $q \ll M_W$, identify Fermi coupling constant

$$G_F = rac{g^2}{8M_W^2}$$

Moving to the nucleon system, we face

$$\langle p|ar{u}\gamma^{\mu}(1-\gamma^{5})d|n
angle$$

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Symmetries to the rescue! CVC & PCAC define new *nucleon* currents

$$V^{\mu} + A^{\mu} pprox g_V(q^2) \gamma^{\mu} (1 - \lambda \gamma^5)$$

where $g_V(q^2) \approx 1$ and λ from the lattice

Strong interaction introduces extra terms into the vertex \rightarrow Construct all Lorentz invariants

$$\langle p | V^{\mu} | n \rangle = \bar{p} \left[g_{V} \gamma^{\mu} + \frac{g_{M} - g_{V}}{2M} \sigma^{\mu\nu} q_{\nu} + i \frac{g_{S}}{2M} q^{\mu} \right] n$$

$$\langle p | A^{\mu} | n \rangle = \bar{p} \left[g_{A} \gamma^{\mu} \gamma^{5} + \frac{g_{T}}{2M} \sigma^{\mu\nu} q_{\nu} \gamma^{5} + i \frac{g_{P}}{2M} q^{\mu} \gamma^{5} \right] n$$

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Introduction of recoil ($\sim q/M$) terms

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Introduction of recoil ($\sim q/M$) terms

CVC requires $g_S = 0 \& g_M = \mu_p^{an} - \mu_n = 4.7$

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Relativistic generalization in Breit frame

$$\langle f|V^0+A^0|i
angle\propto\sum_{LM}(-)^{J_f-M_f}\left(egin{array}{cc} J_f&L&J_i\ -M_f&M&M_i\end{array}
ight)(Y^M_L)^*F_L(q^2)$$

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angle \propto \sum_{LM} (-)^{J_f - M_f} \begin{pmatrix} J_f & L & J_i \\ -M_f & M & M_i \end{pmatrix} (Y_L^M)^* F_L(q^2)$$

Form factors \sim reduced matrix elements

Immediately faced with several issues:

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Here the going gets rough \rightarrow severe approximations

Weak current in strongly bound system?

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Relativistic nuclear wave functions

- \rightarrow Non-relativistic nucleons
 - expand operator to $\mathcal{O}(v/c)$
 - Incomplete wave function basis, core polarization

Standard Model Calculation: Nucleus

Final state interactions

1. Coulomb interaction

Standard Model Calculation: Nucleus

Final state interactions

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 \rightarrow Fermi function, induced Coulomb terms

Final state interactions

1. Coulomb interaction

Make several approximations

• Initial & Final Coulomb potentials are same

Final state interactions

1. Coulomb interaction

Make several approximations

- Initial & Final Coulomb potentials are same
- Typically neglect intermediate decays

Standard Model Calculation: Nucleus

Final state interactions

2. EW Radiative corrections

Standard Model Calculation: Nucleus

Final state interactions

2. EW Radiative corrections

+ higher orders, γW boxes: see previous talks

Must consider total nuclear + atomic Hamiltonian

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Changes

- Available phase space
- Final state interactions
- Opens new decay modes (bound & exchange)

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Require atomic wave functions

- Central & static potential
- Sudden approximation

Similar as atomic system, but changes

- Available phase space
- Molecular excitation, ionization
- Recoil correction & distribution

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Enter quantum chemistry

- Born-Oppenheimer approximation
- MOLCAO

Current status

Large scale gap to cross:

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$$N(W)dW = \frac{G_V^2 V_{ud}^2}{2\pi^3} F_0(Z, W) L_0(Z, W) U(Z, W) R_N(W, W_0, M)$$

$$\times Q(Z, W, M) R(W, W_0) S(Z, W) X(Z, W) r(Z, W)$$

$$\times C(Z, W) D_C(Z, W, \beta_2) D_{FS}(Z, W, \beta_2)$$

$$\times pW(W_0 - W)^2 dW$$

LH et al., Rev. Mod. Phys. 90 (2018) 015008; 1709.07530

80 years of history, in detail

Item	Effect	Formula	Magnitude
1	Phase space factor	$pW(W_0-W)^2$	Unity or larger
2	Traditional Fermi function	F ₀	
3	Finite size of the nucleus	L ₀	
4	Radiative corrections	R	$10^{-1} - 10^{-2}$
5	Shape factor	С	
6	Atomic exchange	X	
7	Atomic mismatch	r	

Added/Improved/Didactic

Item	Effect	Formula	Magnitude
8	Atomic screening	S	
9	Shake-up	See 7	
10	Shake-off	See 7	
11	Isovector correction	C_{I}	
12	Recoil Coulomb correction	Q	10-3 10-4
13	Diffuse nuclear surface	U	10 -10
14	Nuclear deformation	D_{FS}	
15	Recoiling nucleus	R _N	
16	Molecular screening	ΔS_{Mol}	
17	Molecular exchange	Case by case	

Added/Improved/Didactic

Performance summary

Comparison against numerical results for superallowed & mirror transitions

Agreement is very good

Serves as input for several experiments, C++ code available

L. H. et al., 1803.00525, github.com/leenderthayen/BSG

Order of magnitude estimates

Nuclear structure sensitivity in shape factor

$$C(Z,W) \sim 1 \pm \frac{4}{3} \frac{W}{M_N} \frac{\boldsymbol{b}}{Ac} \pm \frac{4\sqrt{2}}{21} \alpha ZWR\boldsymbol{\Lambda} - \frac{1}{3WMc} (\pm 2\boldsymbol{b} + \boldsymbol{d})$$

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Fill in typical numbers to obtain

Matrix element	Name	Slope (% MeV^{-1})
b	Weak Magnetism	0.5
d	Induced Tensor	0.1
Λ	Induced Pseudoscalar	0.1

Weak magnetism is generally more stable than others \rightarrow essential to get this right

Weak magnetism

Mirror nuclei have CVC-determined WM

open: l + 1/2, closed: l - 1/2

Weak magnetism

'Easy' matrix elements only accurate to 10-20%

Weak magnetism

How does shell model perform right now?

 $\Delta b/Ac = 1
ightarrow 0.1\%~{
m MeV^{-1}}$

Still large discrepancies for d/Ac

PHYSICAL REVIEW C 95, 035501 (2017)

2_1^+ to $3_1^+ \gamma$ width in ²²Na and second class currents

S. Triambak,^{1,2,*} L. Phuthu,¹ A. García,³ G. C. Harper,³ J. N. Orce,¹ D. A. Short,³ S. P. R. Steininger,³ A. Diaz Varela,⁴ R. Dunlop,⁴ D. S. Jamieson,⁴ W. A. Richter,¹ G. C. Ball,⁵ P. E. Garrett,⁴ C. E. Svensson,⁴ and C. Wrede^{3,6}

$21(6) \geq d/Ac \geq 3(6)$

Factor 7 differences depending on shell model results \rightarrow killer!

Challenges

At $\mathcal{O}(10^{-3})$, nuclear structure is main culprit

- Nuclear matrix elements only precise to 10-20%
- Generally: large meson exchange corrections on induced currents
- Isospin multiplet decays are way to go: WM from CVC, induced tensor = 0

At $\leq \mathcal{O}(10^{-4})$, everything breaks

At $\leq \mathcal{O}(10^{-4}),$ everything breaks , but not in the same place!

- Low energy: Atomic & Molecular effects (exchange)
- Endpoint: Final state interactions, excitations
- Radiative corrections: higher order, model dependence
- Low Z: recoil corrections to matrix elements
- High Z: everything electromagnetic

Theoretical spectrum is *theoretically* valid to few 10^{-4}

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Nuclear structure generally is main current generation bottleneck

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Further, radiative & recoil corrections become bottleneck even for nuclear-structure-favorable transitions