

The CP Nature of the Higgs Boson: Introduction

M.J. Ramsey-Musolf

U Mass Amherst



AMHERST CENTER FOR FUNDAMENTAL INTERACTIONS

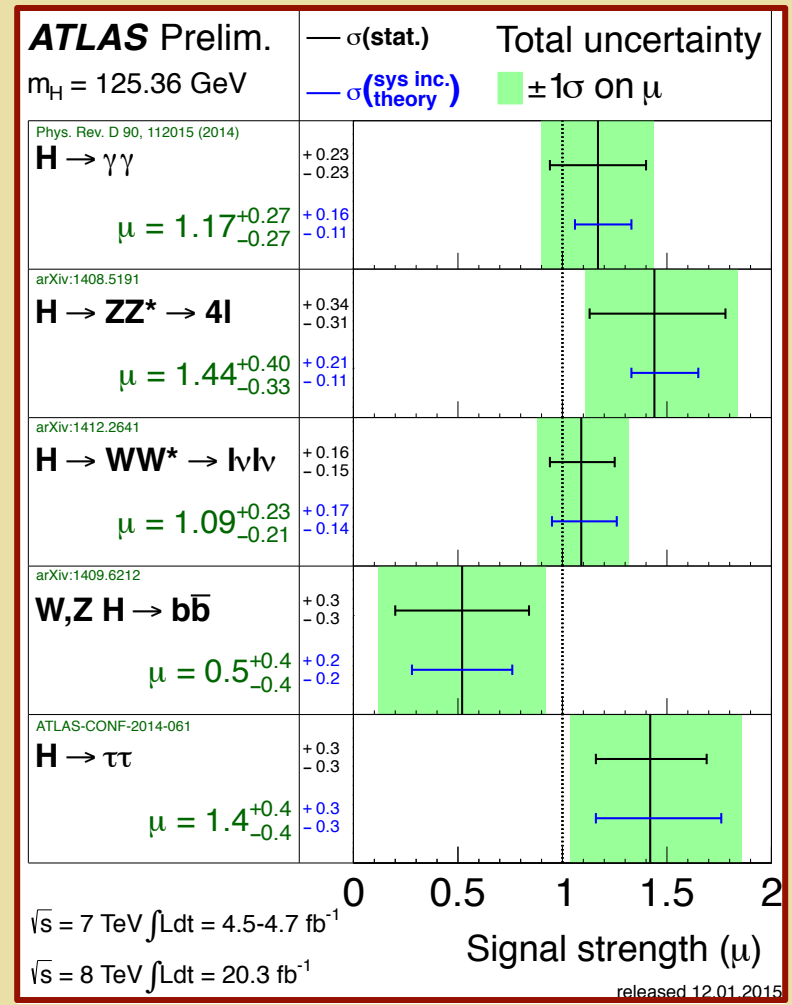
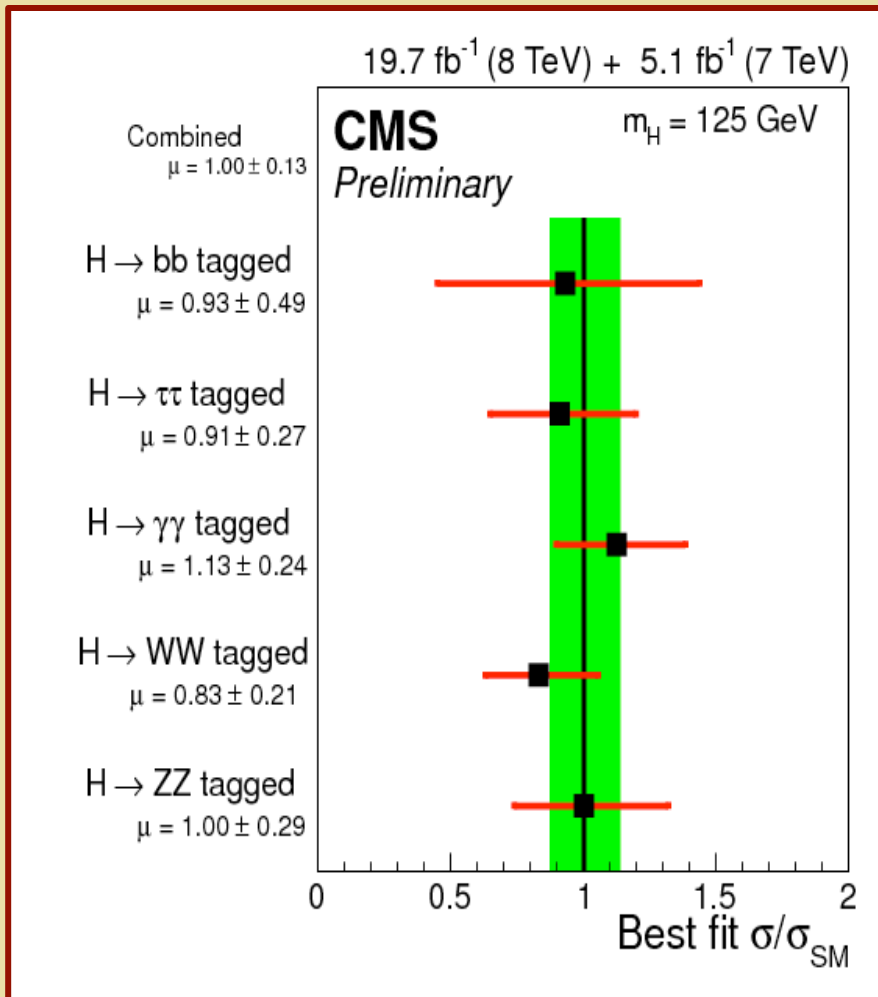
Physics at the interface: Energy, Intensity, and Cosmic frontiers

University of Massachusetts Amherst

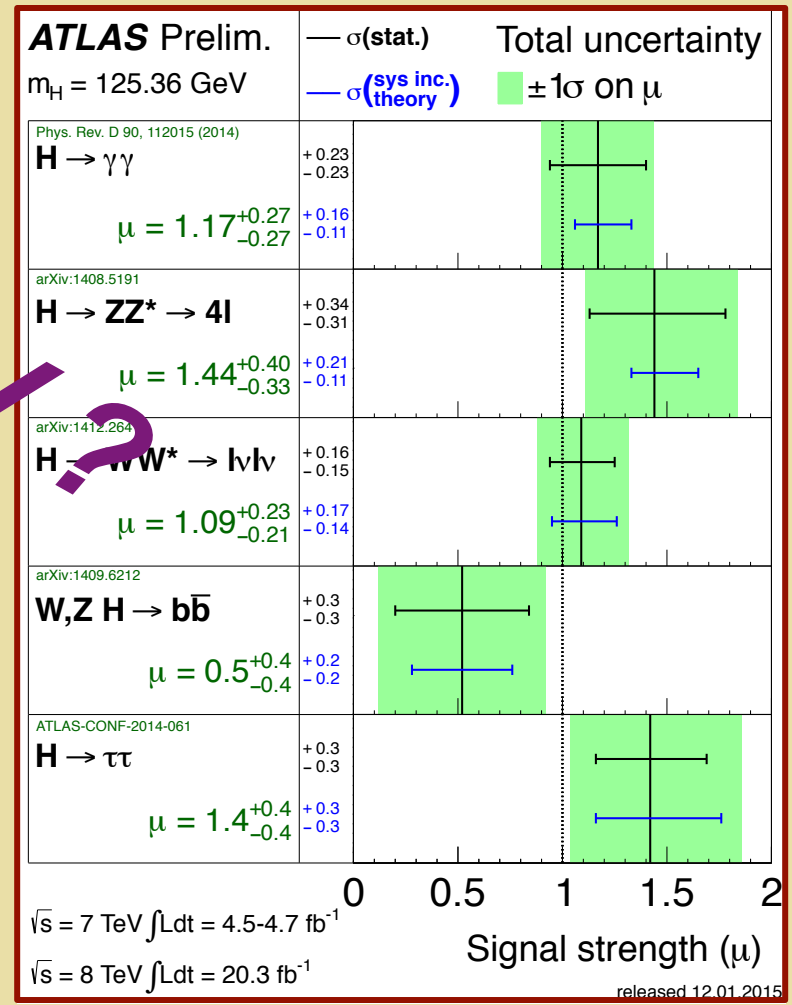
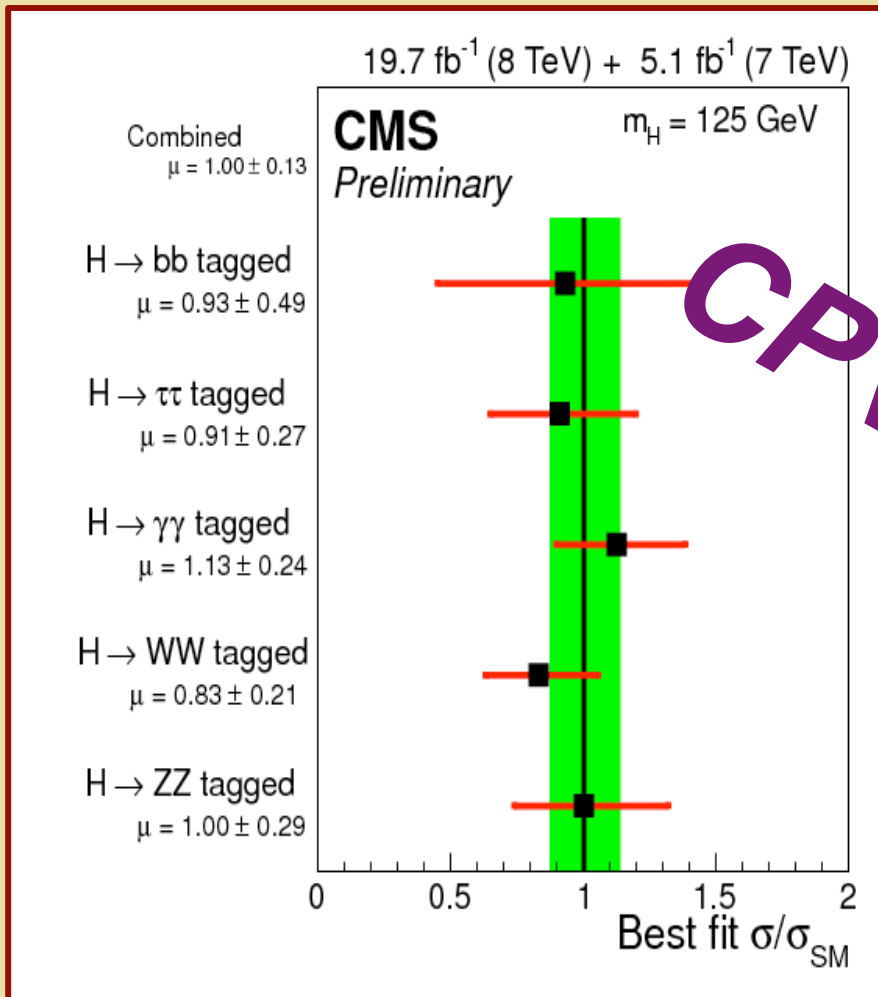
<http://www.physics.umass.edu/acfi/>

ACFI Workshop May 2015

BEH-like Boson Observation



BEH-like Boson Observation



Questions for this Workshop

- What are the sources of CPV in extended Higgs sectors?*
- What are the implications of Higgs sector CPV for determinations of Higgs properties?*
- What are the cosmological implications of Higgs sector CPV*
- What are the potential future collider probes of Higgs sector CPV?*
- What are the “low energy” constraints (present and future) ?*

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Sources of Higgs Sector CPV

CPV & 2HDM: Type I & II

$\lambda_{6,7} = 0$ for simplicity

$$V = \frac{\lambda_1}{2}(\phi_1^\dagger\phi_1)^2 + \frac{\lambda_2}{2}(\phi_2^\dagger\phi_2)^2 + \lambda_3(\phi_1^\dagger\phi_1)(\phi_2^\dagger\phi_2) + \lambda_4(\phi_1^\dagger\phi_2)(\phi_2^\dagger\phi_1) + \frac{1}{2} \left[\lambda_5(\phi_1^\dagger\phi_2)^2 + \text{h.c.} \right] - \frac{1}{2} \left\{ m_{11}^2(\phi_1^\dagger\phi_1) + \left[m_{12}^2(\phi_1^\dagger\phi_2) + \text{h.c.} \right] + m_{22}^2(\phi_2^\dagger\phi_2) \right\}.$$

$$\begin{aligned} \delta_1 &= \text{Arg} \left[\lambda_5^*(m_{12}^2)^2 \right], \\ \delta_2 &= \text{Arg} \left[\lambda_5^*(m_{12}^2)v_1v_2^* \right] \end{aligned}$$

EWSB

$$\delta_2 \approx \frac{1 - \left| \frac{\lambda_5 v_1 v_2}{m_{12}^2} \right|}{1 - 2 \left| \frac{\lambda_5 v_1 v_2}{m_{12}^2} \right|} \delta_1$$

What other extended Higgs sector CPV scenarios ?

Inoue, R-M, Zhang:
1403.4257

Questions for this Workshop

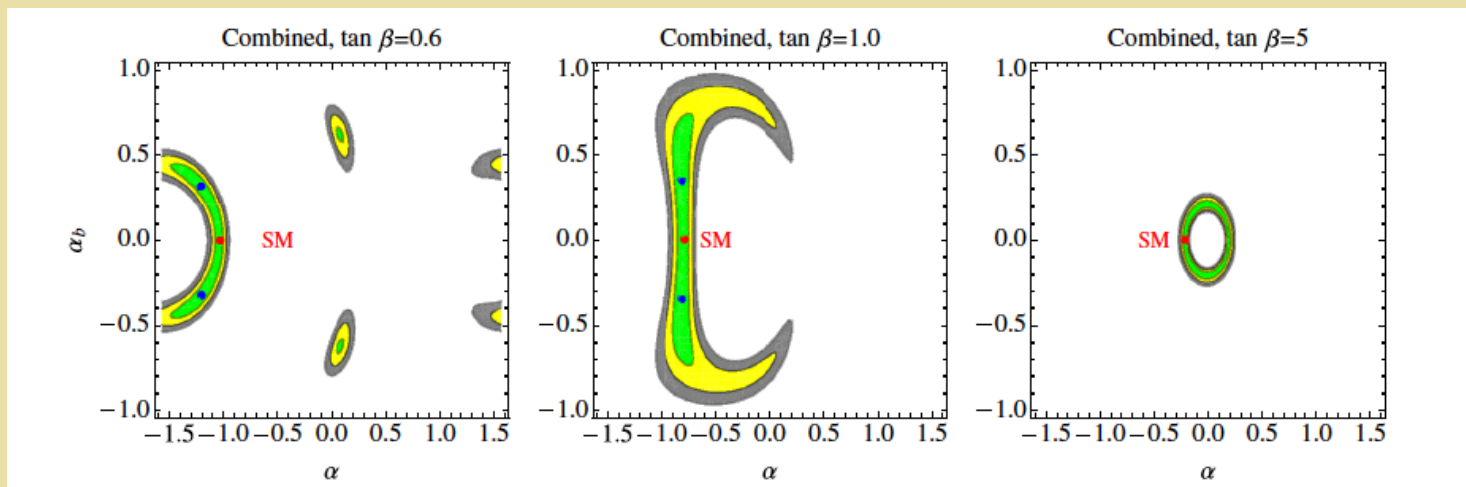
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Implications for Higgs Properties

CPV & 2HDM: Type I & II

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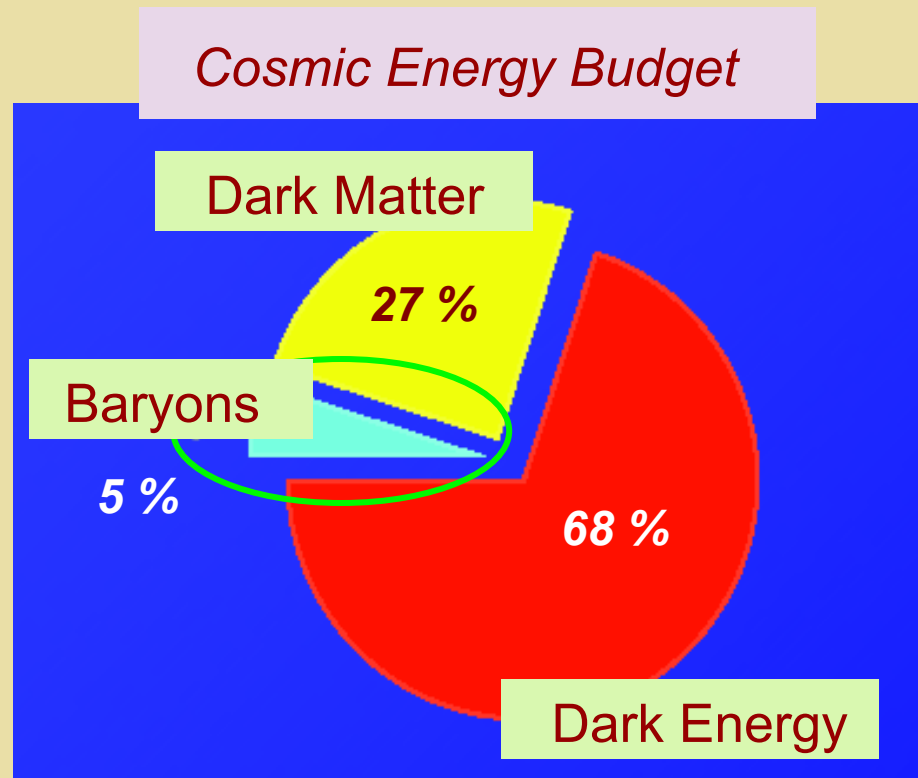


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Cosmological Implications



Baryon asymmetry requires new sources of CPV

Ingredients for Baryogenesis



- *B violation (sphalerons)*
- *C & CP violation*
- *Out-of-equilibrium or CPT violation*

Standard Model

BSM

✓

✓

✗

✓

✗

✓

Ingredients for Baryogenesis



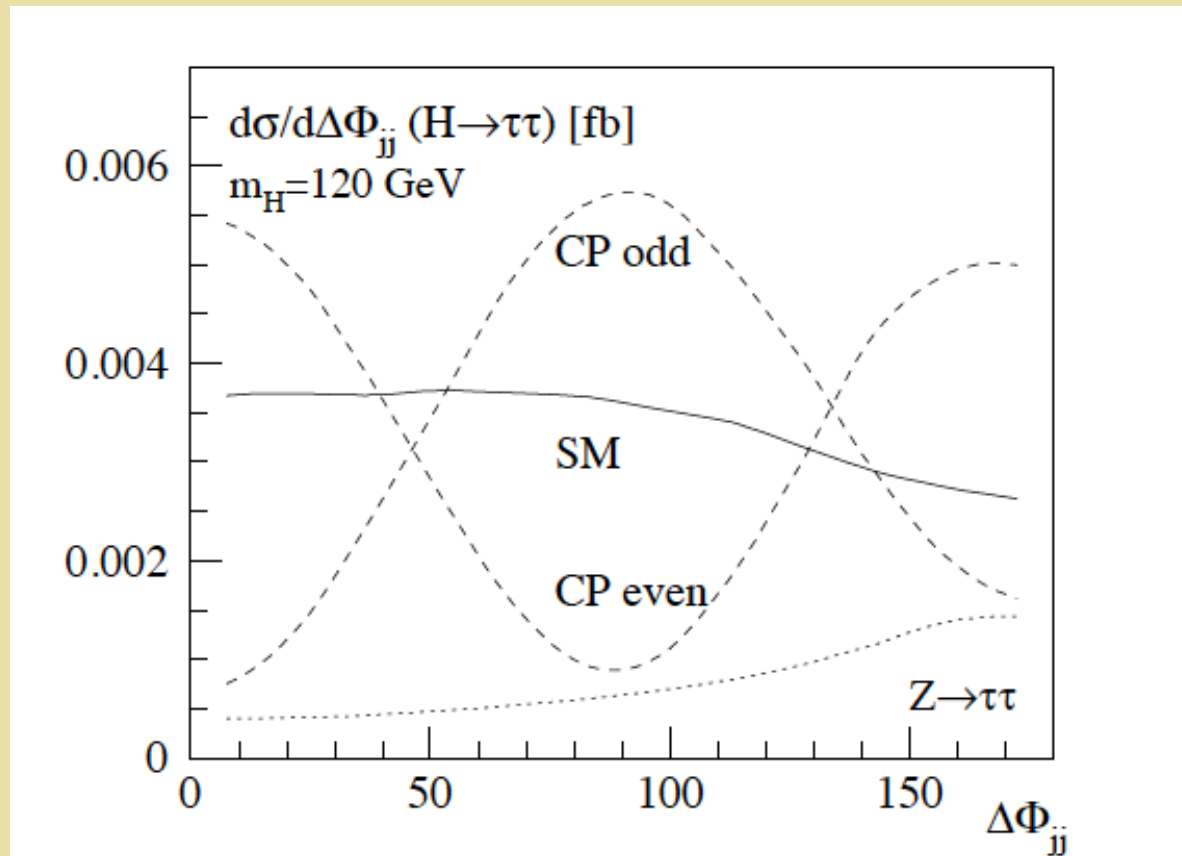
Scenarios: *leptogenesis, EW baryogenesis, Affleck-Dine, asymmetric DM, cold baryogenesis, post-sphaleron baryogenesis...*

	<i>Standard Model</i>	<i>BSM</i>
• <i>B violation (sphalerons)</i>	✓	✓
• <i>C & CP violation</i>	✗	✓
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Collider Probes



WBF: dijet azimuthal distribution

*Plehn, Rainwater, Zeppenfeld,
hep-ph/0105325*

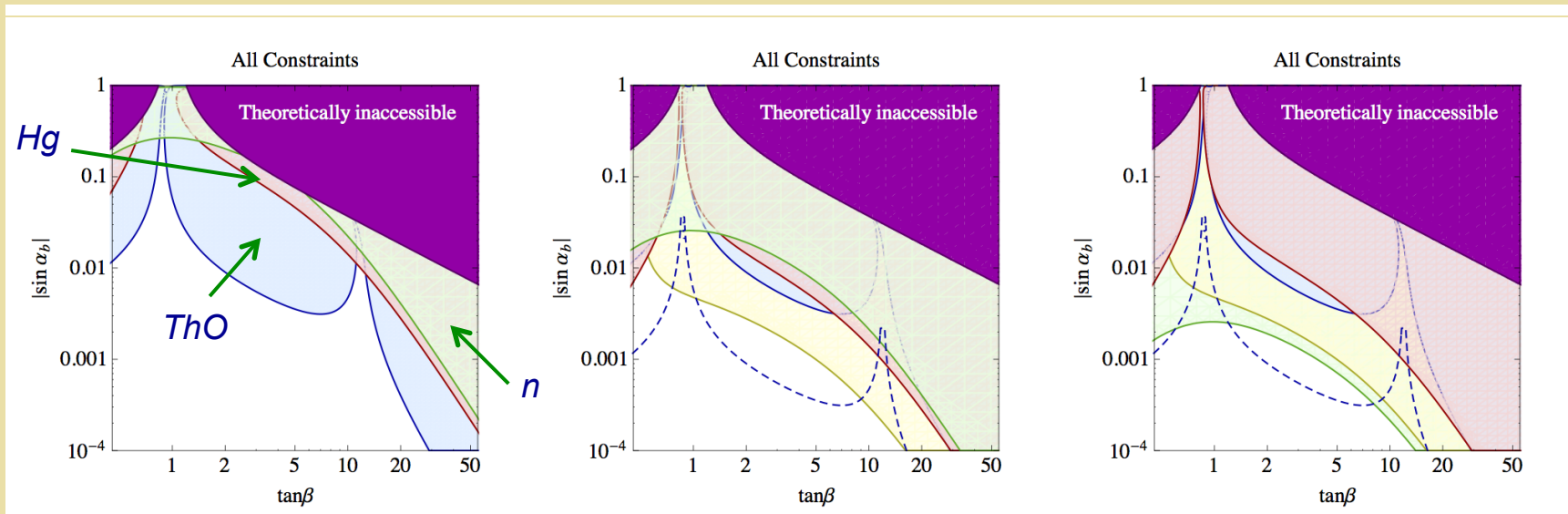
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Low Energy Probes: EDMs

CPV & 2HDM: Type II illustration

$\Lambda_{6,7} = 0$ for simplicity



Present

$\sin \alpha_b$: CPV
scalar mixing

Future:

$d_n \times 0.1$
 $d_A(\text{Hg}) \times 0.1$
 $d_{\text{ThO}} \times 0.1$
 $d_A(\text{Ra})$

Future:

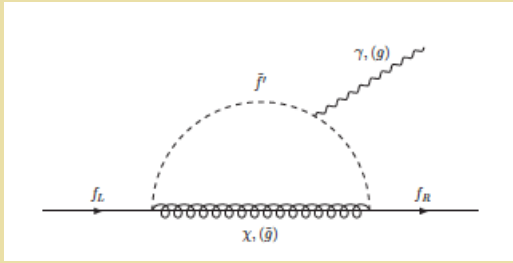
$d_n \times 0.01$
 $d_A(\text{Hg}) \times 0.1$
 $d_{\text{ThO}} \times 0.1$
 $d_A(\text{Ra})$

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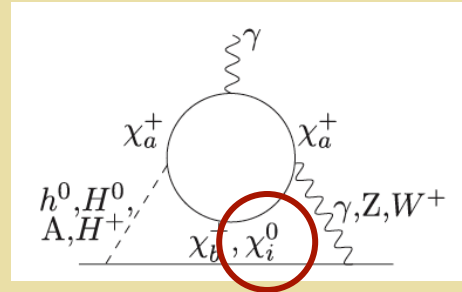
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Putting it all together

EDM Probes: EWB Implications



Heavy sfermions: LHC consistent & suppress 1-loop EDMs



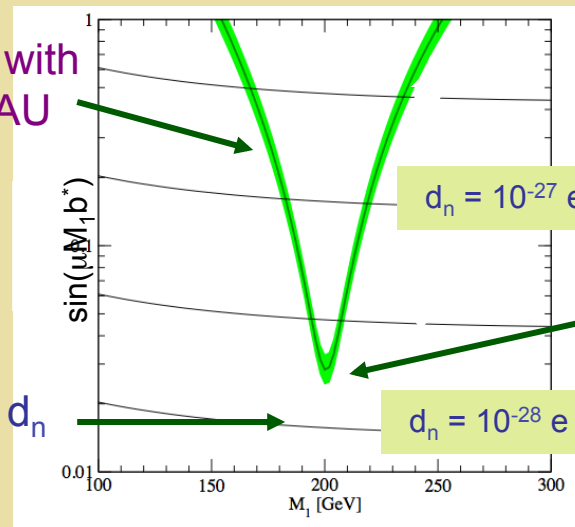
Sub-TeV EW-inos: LHC & EWB - viable but non-universal phases

Viable EWB & CPV:

- EDMs are 2-loop
- CPV is flavor non-diag

Compatible with observed BAU

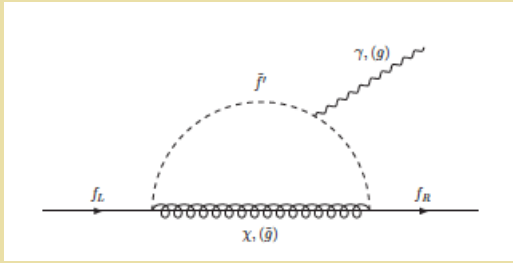
Next gen d_n



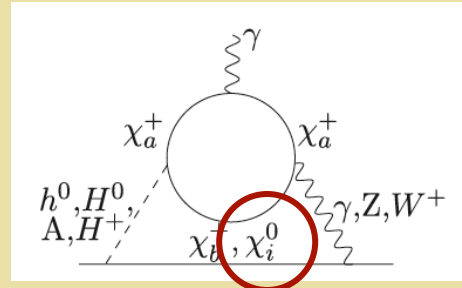
Compressed spectrum (stealthy SUSY)

Li, Profumo, RM '09-' 10

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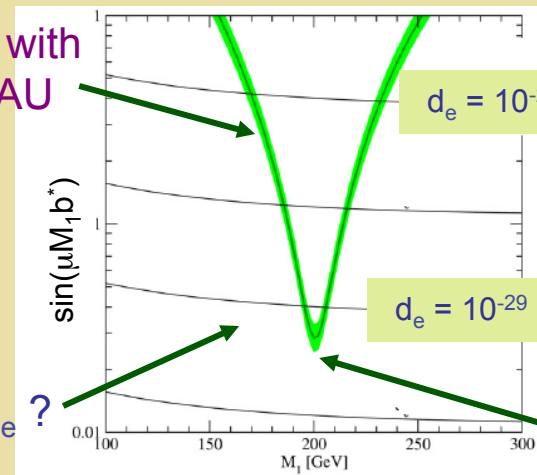


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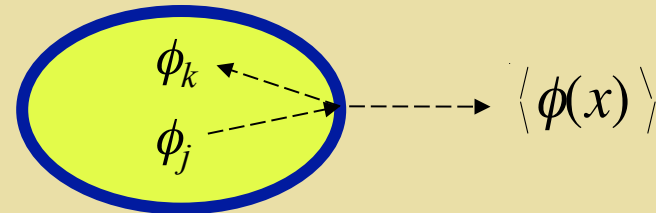
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Scalar asym $\rightarrow n_L$ via Yukawa int