Conformal Tensors in Lovelock Gravity

Lovelock gravity shares important features with Einstein gravity...

But exactly which features?

Do we know them all?

The Riemann-Lovelock Curvature Tensor

David Kastor (Massachusetts U., Amherst). Feb 2012. 12 pp. Published in Class.Quant.Grav. 29 (2012) 155007 DOI: <u>10.1088/0264-9381/29/15/155007</u> e-Print: <u>arXiv:1202.5287</u> [hep-th] | <u>PDF</u>

Conformal Tensors via Lovelock Gravity

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Formulate some basic questions...

- a) Higher Curvature Bianchi Identities
- b) Analogues of 3D GR
- 1. Intro to Lovelock
- 2. Riemann-Lovelock Tensor & "Lovelock Flatness"
- 3. Weyl-Lovelock et. al.
- 4. Further (interesting?) questions

In abundance...

Answers supplied by new higher curvature constructs

What about "conformal Lovelock flatness"?

1) Introduction to Lovelock

$$S = \int d^{D}x \sqrt{-g} \sum_{k=0}^{k_{D}} a_{k} \mathcal{R}^{(k)}$$
Higher curvature analogues of scalar curvature

$$\mathcal{R}^{(k)} = \delta_{b_{1}...b_{2k}}^{a_{1}...a_{2k}} R_{a_{1}a_{2}}^{b_{1}b_{2}} \dots R_{a_{2k-1}a_{2k}}^{b_{2k-1}b_{2k}}$$

$$\delta_{a_{1}...a_{n}}^{b_{1}...b_{n}} = \delta_{[a_{1}}^{b_{1}} \cdots \delta_{a_{n}]}^{b_{n}}$$

$$\mathcal{R}^{(0)} = 1$$
Cosmological constant term

$$\mathcal{R}^{(1)} = R$$
Einstein-Hilbert term

$$\mathcal{R}^{(2)} = \frac{1}{6} \left(R_{ae}{}^{cd}R_{cd}{}^{be} - 2R_{ad}{}^{bc}R_{c}{}^{d} - 2R_{c}{}^{b}R_{a}{}^{c} + R_{a}{}^{b}R \right)$$
Gauss-Bonnet term

$$\mathcal{R}^{(k)}$$
Euler density in D=2k dimensions
- vanishes for D<2k
- variation vanishes in D=2k
$$k_{D} = \left[\frac{D-1}{2} \right]$$

The nice thing about Lovelock...

$$S = \int d^D x \sqrt{-g} \sum_{k=0}^{k_D} a_k \mathcal{R}^{(k)}$$

$$\sum_{k=0}^{k_D} a_k \mathcal{G}^{(k)a}{}_b = 0$$

Equations of motion depend only on Riemann tensor and not its derivatives

- no 4th derivatives of metric
- same initial data as GR

- no ghosts

Higher curvature analogues of Einstein tensor

$$\begin{split} \mathcal{G}_{a}^{(k)b} &= \frac{(2k+1)\alpha_{k}}{2} \, \delta_{ad_{1}\dots d_{2k}}^{bc_{1}\dots c_{2k}} \, R_{c_{1}c_{2}}^{d_{1}d_{2}} \, \dots \, R_{c_{2k-1}c_{2k}}^{d_{2k-1}d_{2k}} \\ \nabla_{a}\mathcal{G}^{(k)a}{}_{b} &= 0 \qquad \text{Covariantly conserved} \\ \mathcal{G}^{(1)a}{}_{b} & \text{Einstein tensor} \\ \mathcal{G}^{(k)a}{}_{b} & \text{Vanishes for D < 2k+1} \end{split}$$

A)
$$\nabla_a \mathcal{G}^{(k)a}{}_b = 0$$
 Covariantly conserved

k=1 Follows from twice contracted Bianchi identity

$$\nabla_{[a}R_{bc]}{}^{de} = 0 \qquad \Longrightarrow \qquad 0 = \nabla_{[a}R_{bc]}{}^{bc}$$
$$= \frac{1}{3}(\nabla_{a}R - 2\nabla_{b}R^{b}{}_{a})$$
$$= -\frac{2}{3}\nabla_{b}G^{b}{}_{a}$$

Is there an analogue of the uncontracted Bianchi identity for k>1?

Is there a higher curvature Lovelock analogue of the Riemann tensor in this sense?

B) Vacuum GR in D=3

$$G_{ab} = 0 \implies R_{ab}{}^{cd} = 0$$

All solutions to Einstein's equation are flat

Both Riemann and Ricci tensors have 6 independent components

3 x 3 symmetric tensors

Or simple Lovelock-type construction ...

$$\delta_{abgh}^{cdef} R_{ef}{}^{gh} = \frac{1}{6} \left(R_{ab}{}^{cd} - 4\delta_{[a}^{[c}R_{b]}{}^{d]} + \delta_{ab}^{cd} R \right) \qquad \begin{array}{l} \text{Relation is true in} \\ \text{all dimensions} \end{array}$$

LHS vanishes in D=3, determining Riemann tensor in terms of its contractions

B) Vacuum GR in D=3

$$G_{ab} = 0 \implies R_{ab}{}^{cd} = 0$$
 All solutions to Einstein's equation are flat
Is there an analogue of this for k>1?
 $\mathcal{R}^{(k)}$ Euler density in D=2k dimensions \implies Look at "pure" kth order Lovelock gravity in D=2k+1
 $S_{pure} = \int d^D x \sqrt{-g} \mathcal{R}^{(k)}$ Only kth order Lovelock Trivial in D=2k, like GR in D=2
D=2k+1 \implies This is highest order Lovelock will asymptote to solutions of pure kth order theory in high curvature regime

B) Vacuum GR in D=3

$$G_{ab} = 0 \implies R_{ab}{}^{cd} = 0$$

Is there an analogue of this for k>1?

 $\mathcal{R}^{(k)}$ Euler density in D=2k dimensions



Look at "pure" kth order Lovelock gravity in D=2k+1

$$S_{pure} = \int d^D x \sqrt{-g} \mathcal{R}^{(k)}$$

Only kth order Lovelock term in action

Is there a higher curvature Lovelock flatness condition, such that all solutions to pure k^{th} order Lovelock in D=2k+1 are k^{th} order Lovelock flat?

Like question 1, this calls for a higher curvature analogue of the Riemann tensor

2) Riemann-Lovelock tensors & "Lovelock flatness" Call this Riemann^(k) tensor

$$\mathcal{R}^{(k)}_{a_1b_1\dots a_kb_k}{}^{c_1d_1\dots c_kd_k} \equiv R_{[a_1b_1}{}^{[c_1d_1}R_{a_2b_2}{}^{c_2d_2}\cdots R_{a_kb_k]}{}^{c_kd_k]}$$

Tensor of type (2k,2k), vanishes for D<2k and satisfies... Like all 1D spaces are k=1 Lovelock flat

$$\mathcal{R}_{a_{1}...a_{2k}b_{1}...b_{2k}}^{(k)} = \mathcal{R}_{[a_{1}...a_{2k}]b_{1}...b_{2k}}^{(k)} = \mathcal{R}_{a_{1}...a_{2k}[b_{1}...b_{2k}]}^{(k)} = \mathcal{R}_{b_{1}...b_{2k}a_{1}...a_{2k}}^{(k)}$$

$$\mathcal{R}_{[a_{1}...a_{2k}b_{1}]}^{(k)} \stackrel{b_{2}...b_{2k}}{=} 0$$
Bianchi identities
Symmetries

Analogous to familiar properties of Riemann tensor

$$\kappa^{\text{th}}$$
 order Lovelock flatness $\mathcal{R}^{(k)}_{a_1b_1...a_kb_k} c_1 d_1...c_k d_k = 0$ or Riemann^(k) flat

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Taking traces...

$$\mathcal{R}^{(k)} = \mathcal{R}^{(k)}_{a_1...a_{2k}}{}^{a_1...a_{2k}}$$
$$\mathcal{R}^{(k)b}_a = \mathcal{R}^{(k)}_{ac_1...c_{2k-1}}{}^{bc_1...c_{2k-1}}$$

Tracing over all pairs of indices gives back scalar Lovelock interaction terms

Ricci^(k) tensor is an analogue of Ricci tensor

$$\mathcal{G}_a^{(k)b} = k\mathcal{R}_a^{(k)b} - (1/2)\delta_b^a \mathcal{R}^{(k)}$$

Einstein^(k) tensor appears in Lovelock equation of motion

$$0 = \nabla_{[a} \mathcal{R}_{b_1 \dots b_{2k}]}^{(k)} {}^{b_1 \dots b_{2k}}$$
$$= \frac{1}{2k+1} \left(\nabla_a \mathcal{R}^{(k)} - 2k \nabla_b \mathcal{R}^{(k)b}{}_a \right)$$
$$= -\frac{2}{2k+1} \nabla_b \mathcal{G}^{(k)b}{}_a$$

Fully contracted Bianchi identity yields vanishing divergrance for Einstein^(k) tensors

Answers 1st question

Demonstrates some relevance for Riemann-Lovelock tensors

Analogue of Pure kth order Lovelock in D=2k+1 vacuum GR in D=3 $S_{pure} = \int d^{2k+1}x \sqrt{-g} \mathcal{R}^{(k)}$ $\mathcal{G}_a^{(k)b} = k\mathcal{R}_a^{(k)b} - (1/2)\delta_b^a \mathcal{R}^{(k)} = 0$ Yes 2nd Question Are all solutions kth order Lovelock flat? $\mathcal{R}^{(k)}_{a_1b_1\dots a_kb_k}{}^{c_1d_1\dots c_kd_k} \equiv R_{[a_1b_1}{}^{[c_1d_1}R_{a_2b_2}{}^{c_2d_2}\cdots R_{a_kb_k]}{}^{c_kd_k]}$ Same number of independent components D=2k+1 as symmetric (2k+1)x(2k+1) tensor $\mathcal{R}_{a}^{(k)b} = 0 \implies \mathcal{R}_{a_{1}b_{1}\dots a_{k}b_{k}}^{(k)} c_{1}d_{1}\dots c_{k}d_{k} = 0$ Can show

Are there interesting spacetimes that are higher order Lovelock flat, but not Riemann flat?

Large set of examples...

Riemann^(k) tensor vanishes for any spacetime of dimension D < 2k

Can build higher dimensional Riemann^(k) flat spacetimes by adding flat directions

Interesting example in D=2k+1...

Static, spherically symmetric solutions of pure kth order Lovelock are missing solid angle spacetimes

$$ds_{2k+1}^2 = -dt^2 + dr^2 + \alpha^2 r^2 d\Omega_{2k-1}^2$$

Static, spherically symmetric solutions of pure kth order Lovelock are missing solid angle spacetimes

$$ds_{2k+1}^2 = -dt^2 + dr^2 + \alpha^2 r^2 d\Omega_{2k-1}^2$$

$$R_{\mu\nu}^{\ \rho\sigma} = \frac{2}{\alpha^2 r^2} (1 - \alpha^2) \delta_{\mu\nu}^{\rho\sigma} \qquad \mu, \nu = 1, \dots, 2k-1 \quad \begin{array}{l} \text{Angular} \\ \text{coordinates} \\ \text{on sphere} \end{array}$$

Only nonzero curvature components

Curved for $\, \alpha
eq 1 \,$

Riemann^(k) tensor

$$\mathcal{R}_{a_1b_1\dots a_kb_k}^{(k)}{}^{c_1d_1\dots c_kd_k} \equiv R_{[a_1b_1}{}^{[c_1d_1}R_{a_2b_2}{}^{c_2d_2}\cdots R_{a_kb_k]}{}^{c_kd_k]} = 0$$

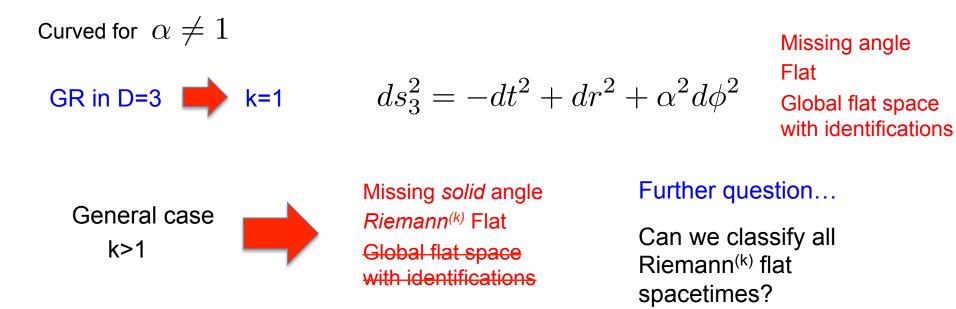
Involves anti-symmetrization over 2k indices, but only 2k-1 are available...

Static, spherically symmetric solutions of pure kth order Lovelock are missing solid angle spacetimes

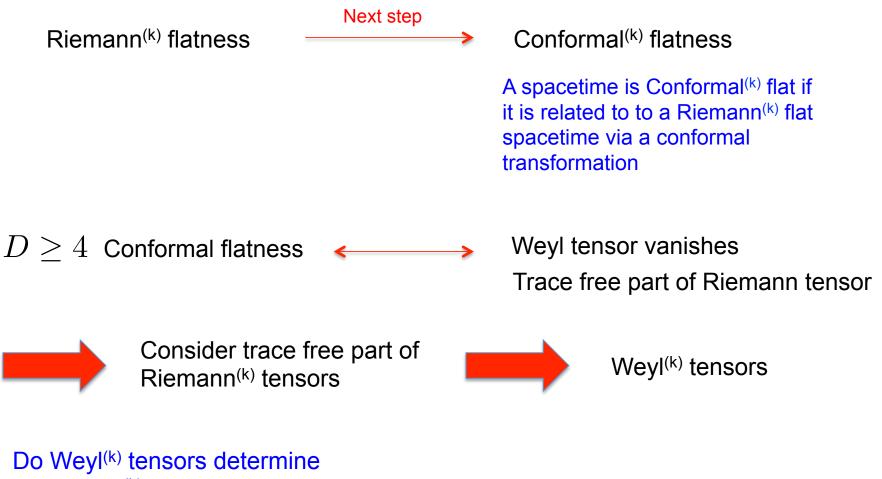
$$ds_{2k+1}^2 = -dt^2 + dr^2 + \alpha^2 r^2 d\Omega_{2k-1}^2$$

$$R_{\mu\nu}^{\ \rho\sigma} = \frac{2}{\alpha^2 r^2} (1 - \alpha^2) \delta_{\mu\nu}^{\rho\sigma} \qquad \mu, \nu = 1, \dots, 2k-1 \quad \begin{array}{l} \text{Angular} \\ \text{coordinates} \\ \text{on sphere} \end{array}$$

Only nonzero curvature components



Conformal tensors in Lovelock...



Conformal^(k) flatness?

First recall some other constructs...

$$W_{ab}{}^{cd} = R_{ab}{}^{cd} - 4\delta^{[c}_{[a}S_{b]}{}^{d]} \qquad \text{Weyl tensor}$$

$$S_{a}{}^{b} = \frac{1}{D-2} \left(R_{a}{}^{b} - \frac{1}{2(D-1)} \delta^{b}_{a}R \right) \qquad \text{Schouten tensor}$$

$$C_{ab}{}^{c} = 2\nabla_{[a}S_{b]}{}^{c}$$

$$C_{ab}{}^{b} = 0 \qquad \text{Cotton tensor}$$

Conformal transformations

$$\tilde{g}_{ab} = e^{2f} g_{ab}$$

$$\tilde{W}_{ab}{}^{cd} = e^{-2f} W_{ab}{}^{cd}$$
$$\tilde{C}_{ab}{}^{c} = e^{-2f} \left(C_{ab}{}^{c} - W_{ab}{}^{cd} \nabla_d f \right)$$

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 $\mathbf{O} \mathbf{f}$ Conformal transformations

$$\tilde{g}_{ab} = e^{2J} g_{ab}$$

$$\tilde{W}_{ab}{}^{cd} = e^{-2f} W_{ab}{}^{cd}$$

$$\tilde{C}_{ab}{}^{c} = e^{-2f} \left(C_{ab}{}^{c} - W_{ab}{}^{cd} \nabla_{d} f \right)$$

$$D=3 \Rightarrow W_{ab}{}^{cd} = 0$$

Vanishes in D=3

$$\bigvee_{\substack{\mathbf{v}\\abgh}} V_{ef}^{gh} = (1/6) W_{ab}^{cd}$$

Cotton tensor is conformally invariant

Conformal flatness $C_{ab}{}^c = 0$ condition in D=3

Weyl tensor not defined

All metrics are locally conformally flat

Conformal tensors in Lovelock...

Define Weyl^(k) tensor as traceless part of Riemann^(k) tensor

$$\mathcal{W}_{a_1\dots a_{2k}}^{(k)} \overset{b_1\dots b_{2k}}{=} = \mathcal{R}_{a_1\dots a_{2k}}^{(k)} \overset{b_1\dots b_{2k}}{=} + \sum_{p=1}^{2k} \alpha_p \, \delta_{[a_1\dots a_p}^{[b_1\dots b_p} \mathcal{R}_{a_{p+1}\dots a_{2k}]}^{(k)} \overset{b_{p+1}\dots b_{2k}]}{=} \\ \alpha_p = \left(\frac{(2k)!}{(2k-p)!}\right)^2 \, \frac{(-1)^p (D - (4k-1))!}{p! (D - (4k-p-1))!}$$



Riemann^(k) tensor determined by its traces for D<4k Expect Weyl^(k) tensor is nontrivial only for $D \ge 4k$

- D < 4k 1 Weyl^(k) tensor undefined because of divergent coefficients Like Weyl tensor in D=1,2
- D = 4k 1 Weyl^(k) tensor defined, but vanishes identically Like Weyl tensor in D=3

Schouten^(k) and Cotton^(k) tensors

$$\mathcal{W}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k}} = \mathcal{R}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k}} - (2k)^2 \delta^{[b_1}_{[a_1} \,\mathcal{S}^{(k)}_{a_2\dots a_{2k}]}{}^{b_2\dots b_{2k}]}$$

$$\mathcal{C}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k-1}} = 2k \,\nabla_{[a_1} \mathcal{S}_{a_2\dots a_{2k}]}^{(k)}{}^{b_1\dots b_{2k-1}}$$

All in parallel with k=1 case....

 $\mathcal{C}_{a_1\dots a_{2k-1}c}^{(k)}{}^{b_1\dots b_{2k-2}c} = 0 \qquad \text{Traceless}$

$$\nabla_c \mathcal{W}_{a_1...a_{2k}}^{(k)} {}^{cb_1...b_{2k-1}} = (D - (4k - 1)) \mathcal{C}_{a_1...a_{2k}}^{(k)} {}^{b_1...b_{2k-1}}$$

Conformal transformations
$$\tilde{g}_{ab} = e^{2f}g_{ab}$$

$$\tilde{\mathcal{W}}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k}} = e^{-2kf} \,\mathcal{W}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k}}$$
$$\tilde{\mathcal{C}}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k-1}} = e^{-2kf} \left(\mathcal{C}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k-1}} - \mathcal{W}_{a_1\dots a_{2k}}^{(k)}{}^{b_1\dots b_{2k-1}c} \nabla_c f \right)$$

 $D=4k-1 \longrightarrow Weyl^{(k)}$ tensor vanishes $\longrightarrow Cotton^{(k) is}$ conformally invariant

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Math. Ann. 199, 175–204 (1972) © by Springer-Verlag 1972

On the Bianchi Identities

Ravindra S. Kulkarni

No connection to Lovelock, but roughly the same time period

- Demonstrates properties of Riemann^(k) tensors
- Defines Weyl^(k) tensors and shows conformal invariance

JOURNAL OF MATHEMATICAL PHYSICS VOLUME 12, NUMBER 3 MARCH 1971

The Einstein Tensor and Its Generalizations*

DAVID LOVELOCK Department of Applied Mathematics, University of Waterloo, Waterloo, Ontario, Canada

(Received 27 August 1970)

Conformal transformation of Weyl tensor...

Let $A_{ab}{}^{cd}$ satisfy $A_{abcd} = A_{[ab]cd} = A_{ab[cd]} = A_{cdab}$

Traces $A_a{}^c = A_{ab}{}^{cb}$ $A = A_a{}^a$

Trace free part

$$A_{ab}^{(t)\,cd} = A_{ab}^{\ cd} - \frac{4}{D-2}\delta_{[a}^{[c}A_{b]}^{\ d]} + \frac{2}{(D-1)(D-2)}\delta_{ab}^{cd}A$$

Let
$$\tilde{A}_{ab}{}^{cd} = A_{ab}{}^{cd} + \delta^{[c}_{[a}\Lambda_{b]}{}^{d]}$$
 with $\Lambda_{ab} = \Lambda_{ba}$

Can show that....

$$\tilde{A}_{ab}^{(t)\,cd} = A_{ab}^{(t)\,cd}$$

Conformal transformation $\tilde{R}_{ab}{}^{cd} = e^{-2f} \left(R_{ab}{}^{cd} + \delta^{[c}_{[a}\Lambda_{b]}{}^{d]} \right)$ $\Lambda_{a}{}^{b} = 4\nabla_{a}\nabla^{b}f + 4(\nabla_{a}f)\nabla^{b}f - 2\delta^{b}_{a}(\nabla_{c}f)\nabla^{c}f$ Analogous construction works for all k



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Conformal^(k) flatness conjectures

k=1 result

D < 2k	Riemann ^(k) tensor vanishes All spacetimes conformal ^(k) flat	No curvature in D=1
D = 2k	Riemann ^(k) tensor has a single component All spacetimes (locally) conformal ^(k) flat?	All D=2 spacetimes are (locally) conformally flat
2k < D < 4k - 1	Weyl ^(k) & Cotton ^(k) tensors not defined All spacetimes (locally) conformal ^(k) flat??	2 < D < 3 No k=1 analogue
		D=3 spacetime is

D = 4k - 1 Conformal^(k) flat if Cotton^(k) tensor vanishes?

 $D \ge 4k$ Conformal^(k) flat if Weyl^(k) tensor vanishes?

D=3 spacetime is conformally flat if Cotton tensor vanishes

 $D \ge 4$ Weyl tensor vanishing implies conformal

flatness

New gravity models?

Recall that low dimensional gravity models make use of conformal tensors...

D=3 Topologically massive gravity (Deser, Jackiw & Templeton – 1982)

Cotton tensor appears in equation of motion

D=3 New massive gravity (Bergshoeff, Hohm & Townsend – 2009)

Schouten tensor is ingredient in action

Perhaps conformal^(k) tensors can be useful in model building associated with Lovelock theories in low(ish) dimensions...

Simple example \longrightarrow Conformal^(k) gravity in D=4k

Recall...

Conformal gravity in D=4
$$S = \int d^4x \sqrt{-g} W_{ab}{}^{cd} W_{cd}{}^{ab}$$

Equation of motion $B_{ab} = 0$

Bach tensor

$$B_a{}^b = (\nabla^d \nabla_c + \frac{1}{2} R_c{}^d) W_{ad}{}^{bc}$$
$$\tilde{B}_a{}^b = e^{-4f} B_a{}^b$$

Symmetric, traceless

Equations of motion are conformally invariant

All Einstein metrics have vanishing Bach tensor

All conformally Einstein spacetimes are solutions to conformal gravity

Simple example \longrightarrow Conformal^(k) gravity in D=4k

$$S = \int d^{4k} x \sqrt{-g} \, \mathcal{W}_{a_1 \dots a_{2k}}^{(k)}{}^{b_1 \dots b_{2k}} \, \mathcal{W}_{b_1 \dots b_{2k}}^{(k)}{}^{a_1 \dots a_{2k}}.$$

Equation of motion

Compare with...

$$B_a{}^b = (\nabla^d \nabla_c + \frac{1}{2} R_c{}^d) W_{ad}{}^{bc}$$

Bach tensor

$$\mathcal{B}_{a}^{(k)b} = \left(\mathcal{R}_{c_{1}...c_{2k-2}}^{(k-1)} d_{1}...d_{2k-2} \nabla^{d_{2k-1}} \nabla_{c_{2k-1}} + \frac{k}{2} \mathcal{R}_{c_{1}...c_{2k-1}}^{(k)} d_{1}...d_{2k-1}\right) \mathcal{W}_{ad_{1}...d_{2k-1}}^{(k)} bc_{1}...c_{2k-1}$$
$$\mathcal{B}_{(ab)}^{(k)} = 0$$

Expect anti-symmetric part of Bach tensor vanishes, but not straightforward to show... As it does for k=1

Also expect Bach^(k) tensor is a conformal invariant, because of conformal invariance of action

Simple example \longrightarrow Conformal^(k) gravity in D=4k

$$S = \int d^{4k} x \sqrt{-g} \, \mathcal{W}_{a_1...a_{2k}}^{(k)} {}^{b_1...b_{2k}} \, \mathcal{W}_{b_1...b_{2k}}^{(k)} {}^{a_1...a_{2k}}.$$
$$\mathcal{B}_a^{(k)b} = \left(\mathcal{R}_{c_1...c_{2k-2}}^{(k-1)} {}^{d_1...d_{2k-2}} \nabla^{d_{2k-1}} \nabla_{c_{2k-1}} + \frac{k}{2} \mathcal{R}_{c_1...c_{2k-1}}^{(k)} {}^{d_1...d_{2k-1}} \right) \mathcal{W}_{ad_1...d_{2k-1}}^{(k)} {}^{bc_1...c_{2k-1}}$$
$$\mathcal{B}_{(ab)}^{(k)} = 0$$

Solved by Einstein^(k) spaces...

$$\mathcal{R}^{(k)}_{a_1\dots a_{2k-1}}{}^{b_1\dots b_{2k-1}} = \alpha \,\delta^{b_1\dots b_{2k-1}}_{a_1\dots a_{2k-1}}$$

Conclusions...

Riemann^(k) tensor looks interesting.

Lots of related questions...