## CPV: EDM's \& Electroweak Baryogenesis

M.J. Ramsey-Musolf<br>U Mass Amherst



## - Amherst Center for Fundamental Interactions

Physics at the interface: Energy, Intensity, and Cosmic frontiers
University of Massachusetts Amherst
http://www.physics.umass.edu/acfi/

## ACFI CPV Workshop

March 2018

## Goals for This Talk

- Illustrate the interplay of EDM searches with EW baryogenesis
- Introduce EDM physics
- $\quad$ Set the stage for remainder of the workshop


## Outline

I. EDM's: The SM \& BSM context
II. The Cosmic Matter-Antimatter Asymmetry
III. Electroweak Baryogenesis: Examples
IV. Outlook

## I. EDMs: The SM \& BSM Context

## EDMs \& SM Physics

$$
d_{n}^{S M} \sim\left(10^{-16} \mathrm{e} \mathrm{~cm}\right) \times \theta_{Q C D}+d_{n}{ }^{C K M}
$$

## EDMs \& SM Physics

$$
\begin{aligned}
& d_{n} \text { SM } \sim\left(10^{-16} \mathrm{e} \mathrm{~cm}\right) \times \theta_{Q C D}+d_{n}^{C K M} \\
& \begin{array}{l}
d_{n}{ }^{C K M}=(1-6) \times 10^{-32} \mathrm{e} \mathrm{~cm} \\
\text { c. Seng arxiv: } 1411.1476
\end{array}
\end{aligned}
$$

## EDMs \& SM Physics

$$
\begin{aligned}
& d_{n} \text { SM } \sim\left(10^{-16} \mathrm{e} \mathrm{~cm}\right) \times \theta_{Q C D}+d_{n}^{C K M} \\
& \begin{array}{l}
d_{n} \text { CKM }=(1-6) \times 10^{-32} \mathrm{e} \mathrm{~cm} \\
\text { c. Seng arxiv: } 1411.1476
\end{array} \\
& * 3.3 \times 10^{-33} \mathrm{e} \mathrm{~cm}<d_{p}<3.3 \times 10^{-32} \mathrm{e} \mathrm{~cm}
\end{aligned}
$$

## EDMs \& BSM Physics

$$
d \sim\left(10^{-16} \mathrm{ecm}\right) \times(v / \Lambda)^{2} \times \sin \phi \times y_{f} F
$$

## EDMs \& BSM Physics

$$
\begin{gathered}
d \sim\left(10^{-16} \mathrm{e} \mathrm{~cm}\right) \times(v / \Lambda)^{2} \times \sin \phi \times y_{f} F \\
\text { CPV Phase: large enough for baryogenesis ? }
\end{gathered}
$$

## EDMs \& BSM Physics

## $d \sim\left(10^{-16} \mathrm{ecm}\right) \times(v / \Lambda)^{2} \times \sin \phi \times y_{f} F$ <br> BSM mass scale: TeV ? Much higher?

## EDMs \& BSM Physics

$$
d \sim\left(10^{-16} \mathrm{e} \mathrm{~cm}\right) \times(v / \Lambda)^{2} \times \sin \phi \times y_{f} F
$$

BSM dynamics: perturbative? Strongly coupled? Dependence on other parameters?

## EDMs \& BSM Physics



## EDMs \& BSM Physics



- Baryon asymmetry
- High energy collisions
- EDMs

Cosmic Frontier
Energy Frontier
Intensity Frontier

## EDMs: New CPV?

| System | Limit (e cm) $^{*}$ | SM CKM CPV | BSM CPV |
| :--- | :--- | :---: | :---: |
| 199 Hg | $7.4 \times 10^{-30}$ | $10^{-33}$ | $10^{-29}$ |
| ThO | $8.7 \times 10^{-29 * *}$ | $10^{-38}$ | $10^{-28}$ |
| n | $3.3 \times 10^{-26}$ | $10^{-31}$ | $10^{-26}$ |

* $95 \% \mathrm{CL} \quad{ }^{* *} \mathrm{e}^{-}$equivalent $\quad$ New $\mathrm{Hf} \mathrm{F}^{+}: 1.3 \times 10^{-28} 1704.07928$


## EDMs: New CPV?

| System | Limit (e cm) $^{*}$ | SM CKM CPV | BSM CPV |
| :--- | :--- | :---: | :---: |
| 199 Hg | $7.4 \times 10^{-30}$ | $10^{-33}$ | $10^{-29}$ |
| ThO | $8.7 \times 10^{-29} * *$ | $10^{-38}$ | $10^{-28}$ |
| n | $3.3 \times 10^{-26}$ | $10^{-31}$ | $10^{-26}$ |

* $95 \%$ CL $\quad$ ** $e^{-}$equivalent $\quad$ New $H f F^{+}: 1.3 \times 10^{-28} 1704.07928$

Mass Scale Sensitivity
$\psi \begin{cases}\operatorname{linn}^{\varphi} \gamma & \sin \phi_{C P} \sim 1 \rightarrow M>5000 \mathrm{GeV} \\ \boldsymbol{e}_{\varphi} & M<500 \mathrm{GeV} \rightarrow \sin \phi_{C P}<10^{-2}\end{cases}$

## EDMs: New CPV?

| System | Limit (e cm) $^{*}$ | SM CKM CPV | BSM CPV |
| :--- | :--- | :---: | :---: |
| 199 Hg | $7.4 \times 10^{-30}$ | $10^{-33}$ | $10^{-29}$ |
| ThO | $8.7 \times 10^{-29 * *}$ | $10^{-38}$ | $10^{-28}$ |
| n | $3.3 \times 10^{-26}$ | $10^{-31}$ | $10^{-26}$ |

* 95\% CL ** e- equivalent New Hf $\mathrm{F}^{+}$: $1.3 \times 10^{-28} 1704.07928$

Not shown: muon

proton
\& nuclei
atoms
~ $100 \times$ better sensitivity

## EDMs: New CPV?

| System | Limit (e cm) $^{*}$ | SM CKM CPV | BSM CPV |
| :--- | :--- | :---: | :---: |
| 199 Hg | $7.4 \times 10^{-30}$ | $10^{-33}$ | $10^{-29}$ |
| ThO | $8.7 \times 10^{-29 * *}$ | $10^{-38}$ | $10^{-28}$ |
| n | $3.3 \times 10^{-26}$ | $10^{-31}$ | $10^{-26}$ |

* $95 \% \mathrm{CL} \quad$ ** $e^{-}$equivalent $\quad$ New $\mathrm{Hf} \mathrm{F}^{+}: 1.3 \times 10^{-28} 1704.07928$

Mass Scale Sensitivity


## EDMs: New CPV?

| System | Limit (e cm) $^{*}$ | SM CKM CPV | BSM CPV |
| :--- | :--- | :---: | :---: |
| 199 Hg | $7.4 \times 10^{-30}$ | $10^{-33}$ | $10^{-29}$ |
| ThO | $8.7 \times 10^{-29} * *$ | $10^{-38}$ | $10^{-28}$ |
| n | $3.3 \times 10^{-26}$ | $10^{-31}$ | $10^{-26}$ |

* $95 \% \mathrm{CL} \quad$ ** $e^{-}$equivalent $\quad$ New $\mathrm{Hf} \mathrm{F}^{+}: 1.3 \times 10^{-28} 1704.07928$

Mass Scale Sensitivity
Challenge for EWBG

- EDMs arise at > 1 loop
- CPV is flavor non-diagonal
- CPV is "partially secluded"


## EDMs: New CPV?

| System | Limit (e cm) $^{*}$ | SM CKM CPV | BSM CPV |
| :--- | :--- | :---: | :---: |
| 199 Hg | $7.4 \times 10^{-30}$ | $10^{-33}$ | $10^{-29}$ |
| ThO | $8.7 \times 10^{-29 * *}$ | $10^{-38}$ | $10^{-28}$ |
| n | $3.3 \times 10^{-26}$ | $10^{-31}$ | $10^{-26}$ |

* $95 \%$ CL $\quad$ ** e- equivalent $\quad$ New Hf $F^{+}: 1.3 \times 10^{-28} 1704.07928$

Mass Scale Sensitivity
This talk

--- EDMs arise at >1 loop

- CPV is flavor non-diagonal
- CPV is "partially secluded"


## II. The Matter-Antimatter Asymmetry

## Cosmic Baryon Asymmetry

$$
Y_{B}=\frac{n_{B}}{s}=(8.82 \pm 0.23) \times 10^{-11}
$$

## One number $\rightarrow$ BSM Physics

## Cosmic Baryon Asymmetry

$$
Y_{B}=\frac{n_{B}}{s}=(8.82 \pm 0.23) \times 10^{-11}
$$

## One number $\rightarrow$ 似... Explanations



## Cosmic Baryon Asymmetry

$$
Y_{B}=\frac{n_{B}}{s}=(8.82 \pm 0.23) \times 10^{-11}
$$

## One number $\rightarrow$ 似... Explanations

Experiment can help:

- Discover ingredients
- Falsify candidates



## Baryogenesis Scenarios



## Baryogenesis Scenarios



## Baryogenesis Scenarios



Era of EWSB: $t_{\text {univ }} \sim 10 \mathrm{ps}$

## Electroweak Baryogenesis

Was $Y_{B}$ generated in conjunction with electroweak symmetry-breaking?

## III. Electroweak Baryogenesis

## EWBG: Ingredients

- Strong first order EWPT: LHC $\rightarrow$ Excluded for the MSSM $\rightarrow$ Possible w/ extensions (e.g., NMSSM)
- CPV: SUSY: Sources same as in MSSM + possible additional; non-SUSY


## EW Phase Transition: Higgs Portal



Increasing $m_{h}$
$\longleftarrow$ New scalars

$$
\mathcal{O}_{4}=\lambda_{\phi H} \phi^{\dagger} \phi H^{\dagger} H \quad \mathbf{+} \ldots
$$

## EW Phase Transition: Higgs Portal



Increasing $m_{h}$
« New scalars

$$
\mathcal{O}_{4}=\lambda_{\phi H} \phi^{\dagger} \phi H^{\dagger} H
$$

$$
+\ldots
$$



- Renormalizable
- $\phi$ : singlet or charged under $\operatorname{SU}(2)_{L} \times U(1)_{Y}$
- Generic features of full theory (NMSSM, GUTS...)
- More robust vacuum stability
- Novel patterns of SSB


## Higgs Portal: Simple Scalar Extensions

| Extension | DOF | EWPT | DM |
| :--- | :---: | :---: | :---: |
| Real singlet: $Z_{又}$ | $\mathbf{1}$ | $\checkmark$ | $\checkmark$ |
| Real singlet: $Z_{2}$ | $\mathbf{1}$ | $\nearrow$ | $\nearrow$ |
| Complex Singlet | 2 | $\nearrow$ | $\nearrow$ |
| EW Multiplets | $3+$ | $\nearrow$ | $\nearrow$ |

May be low-energy remnants of UV complete theory \& illustrative of generic features

## Higgs Portal: Simple Scalar Extensions

| Extension | DOF | EWPT | DM |
| :--- | :---: | :---: | :---: |
| Real singlet: $\mathrm{Z}_{\mathrm{又}}$ | $\mathbf{1}$ | $\checkmark$ | $\checkmark$ |
| Real singlet: $\mathrm{Z}_{2}$ | $\mathbf{1}$ | $\checkmark$ | $\nearrow$ |
| Complex Singlet | $\mathbf{2}$ | $\nearrow$ | $\nearrow$ |
| EW Multiplets | $3+$ | $\nearrow$ | $\nearrow$ |

May be low-energy remnants of UV complete theory \& illustrative of generic features (NMSSM...)

## EW Phase Transition: Singlets



Increasing $m_{h}$
$\longleftarrow$ New scalars

Real Singlet: $\quad \phi \rightarrow S$
Simplest Extension:
two states $h_{1} \& h_{2}$

Profumo, R-M, Shaugnessy JHEP 0708 (2007) 010


## EW Phase Transition: Singlets



Modified Higgs Self-Coupling


Profumo, R-M, Wainwright, Winslow: 1407.5342; see also Noble \& Perelstein 0711.3018


## EW Phase Transition: Singlets



Modified Higgs Self-Coupling


Profumo, R-M, Wainwright, Winslow: 1407.5342; see also Noble \& Perelstein 0711.3018


## EW Phase Transition: Singlets



Increasing $m_{h}$
$\longleftarrow$ New scalars
Resonant di-Higgs production:


No \& RM, arXiv:1310.6035 : LHC Discovery w/ $100 \mathrm{fb}^{-1}$


## EW Phase Transition: Singlets




## EDMs \& EWBG: MSSM + Singlets



Heavy sfermions: LHC consistent \& suppress
1-loop EDMs


Sub-TeV EW-inos: LHC \& EWB viable but non-universal phases

## EDMs \& EWBG: MSSM + Singlets



Heavy sfermions: LHC consistent \& suppress
1-loop EDMs


Li, Profumo, RM ‘09-’10


Sub-TeV EW-inos: LHC \& EWB viable but non-universal phases


## EDMs \& EWBG: MSSM + Singlets



Heavy sfermions: LHC consistent \& suppress
1-loop EDMs


Sub-TeV EW-inos: LHC \& EWB viable but non-universal phases


## EW Multiplets: Two-Step EWPT



Increasing $m_{h}$

$\longleftarrow$ New scalars

- Step 1: thermal loops
- Step 2: tree-level barrier


## EW Multiplets: Two-Step EWPT



Increasing $m_{h}$
$\longleftarrow$ New scalars

- Step 1: thermal loops
- Step 2: tree-level barrier



## EW Multiplets: Two-Step EWPT



Increasing $m_{h}$
$\longleftarrow$ New scalars

Real Triplet $\quad \Sigma \sim(1,3,0)$
Two-step EWPT \& dark matter


## EW Multiplets: Two-Step EWPT




Increasing $m_{h}$
$\longleftarrow$ New scalars


## Two-Step EW Baryogenesis


$H_{j} \quad$ St'd Model Scalar Sector
$\phi \quad B S M$ Scalar Sector: at least one SU(2) _ non-singlet plus possibly gauge singlets: "partially secluded sector CPV"

BSM CPV in $\phi$ H interactions: baryogenesis during step 1

## Two-Step EW Baryogenesis



## Illustrative Model:

New sector: "Real Triplet" $\quad \Sigma$ Gauge singlet $S$
$H \rightarrow$ Set of "SM" fields: 2 HDM
(SUSY: "TNMSSM", Coriano...)
Two CPV Phases:
$\delta_{\Sigma}$ : Triplet phase
$\delta_{S}$ :
Singlet phase

Inoue, Ovanesyan, R-M: 1508.05404

## Two-Step EW Baryogenesis \& EDMs



## Two-Step EW Baryogenesis

Two cases: (A) $\delta_{S}=0 \quad$ (B) $\delta_{\Sigma}=0$


Inoue, Ovanesyan, R-M: 1508.05404

## What is the CP Nature of the Higgs Boson?

- Interesting possibilities if part of an extended scalar sector
- Two Higgs doublets ?

$$
H \rightarrow H_{1}, H_{2}
$$

- New parameters:

$$
\begin{aligned}
& \left.\tan \beta=<H_{1}>/<H_{2}\right\rangle \\
& \sin \alpha_{b}
\end{aligned}
$$

## What is the CP Nature of the Higgs Boson?

- Interesting possibilities if part of an extended scalar sector
- Two Higgs doublets ?

$$
H \rightarrow H_{1}, H_{2}
$$

- New parameters:

| $\tan \beta=<H_{1}>/<H_{2}>$ |  |
| :--- | :--- |
| $\sin \alpha_{b}$ | CPV : scalar-pseudoscalar <br> mixing from $V\left(H_{1}, H_{2}\right)$ |

## Higgs Portal CPV

CPV \& 2HDM: Type I \& I/


## Higgs Portal CPV: EDMs

CPV \& 2HDM: Type II illustration
$\lambda_{6,7}=0$ for simplicity


Present
$\sin \alpha_{b}: C P V$
scalar mixing

Future:

$$
\begin{aligned}
& d_{n} \times 0.1 \\
& d_{A}(\mathrm{Hg}) \times 0.1 \\
& d_{\text {ThO }} \times 0.1 \\
& d_{A}(R a)\left[10^{-27} \mathrm{e} \mathrm{~cm}\right]
\end{aligned}
$$

Future:

$$
\begin{aligned}
& d_{n} \times 0.01 \\
& d_{A}(H g) \times 0.1 \\
& d_{T h O} \times 0.1 \\
& d_{A}(R a)
\end{aligned}
$$

## Higgs Portal CPV: EDMs

CPV \& 2HDM: Type II illustration
$\lambda_{6,7}=0$ for simplicity


Present
$\sin \alpha_{b}: C P V$
scalar mixing

Future:

$$
\begin{aligned}
& d_{n} \times 0.1 \\
& d_{A}(\mathrm{Hg}) \times 0.1 \\
& d_{\text {Tho }} \times 0.1 \\
& d_{A}\left(\text { Ra) }\left[10^{-27} \mathrm{e} \mathrm{~cm}\right]\right.
\end{aligned}
$$

Future:

$$
\begin{aligned}
& d_{n} \times 0.01 \\
& d_{A}(H g) \times 0.1 \\
& d_{T h O} \times 0.1 \\
& d_{A}(R a)
\end{aligned}
$$

## Low-Energy / High-Energy Interplay

## Higgs Portal CPV: Source for EWBG?

Dorsch et al, 1611.05874


$$
\alpha_{b} \propto \delta_{1}-\delta_{2}
$$

## CPV for EWBG



## IV. Outlook

- Searches for permanent EDMs of atoms, molecules, hadrons and nuclei provide powerful probes of BSM physics at the TeV scale and above and constitute important tests of weak scale baryogenesis
- Studies on complementary systems is essential for first finding and then disentangling new CPV \& testing EWBG
- EWBG remains an important baryogenesis scenario for which definitive tests will likely require next generation EDM \& collider studies**
- Analysis of EDM implications of other baryogenesis scenarios is an important and interesting topic $\rightarrow$ Many interesting discussions during remainder of this WS


## Back Up Slides

## Higgs Portal CPV

$$
\begin{aligned}
V= & \frac{\lambda_{1}}{2}\left(\phi_{1}^{\dagger} \phi_{1}\right)^{2}+\frac{\lambda_{2}}{2}\left(\phi_{2}^{\dagger} \phi_{2}\right)^{2}+\lambda_{3}\left(\phi_{1}^{\dagger} \phi_{1}\right)\left(\phi_{2}^{\dagger} \phi_{2}\right)+\lambda_{4}\left(\phi_{1}^{\dagger} \phi_{2}\right)\left(\phi_{2}^{\dagger} \phi_{1}\right)+\frac{1}{2}\left[\lambda_{5}\left(\phi_{1}^{\dagger} \phi_{2}\right)^{2}+\text { h.c. }\right] \\
& -\frac{1}{2}\left\{m_{11}^{2}\left(\phi_{1}^{\dagger} \phi_{1}\right)+\left[m_{12}^{2}\left(\phi_{1}^{\dagger} \phi_{2}\right)+\text { h.c. }\right]+m_{22}^{2}\left(\phi_{2}^{\dagger} \phi_{2}\right)\right\} .
\end{aligned}
$$

$$
\begin{aligned}
& \delta_{1}=\operatorname{Arg}\left[\lambda_{5}^{*}\left(m_{12}^{2}\right)^{2}\right] \\
& \delta_{2}=\operatorname{Arg}\left[\lambda_{5}^{*}\left(m_{12}^{2}\right) v_{1} v_{2}^{*}\right]
\end{aligned}
$$

EWSB

$$
\delta_{2} \approx \frac{1-\left\lvert\, \frac{\lambda_{5} v_{1} v_{2}}{m_{12}^{2}}\right.}{1-2\left|\frac{\lambda_{5} v_{1} v_{2}}{m_{12}^{2}}\right|} \delta_{1}
$$

$h, H^{0}, A^{0} \rightarrow h_{1,2,3}$
$\left(\begin{array}{ccc}-s_{\alpha} c_{\alpha_{b}} & c_{\alpha} c_{\alpha_{b}} & s_{\alpha_{b}} \\ s_{\alpha} s_{\alpha_{b}} s_{\alpha_{c}}-c_{\alpha} c_{\alpha_{c}} & -s_{\alpha} c_{\alpha_{c}}-c_{\alpha} s_{\alpha_{b}} s_{\alpha_{c}} & c_{\alpha_{b}} s_{\alpha_{c}} \\ s_{\alpha} s_{\alpha_{b}} c_{\alpha_{c}}+c_{\alpha} s_{\alpha_{c}} & s_{\alpha} s_{\alpha_{c}}-c_{\alpha} s_{\alpha_{b}} c_{\alpha_{c}} & c_{\alpha_{b}} c_{\alpha_{c}}\end{array}\right)$


## Higgs Portal CPV

$\lambda_{6,7}=0$ for simplicity

$$
0,7 \text { 0, }
$$

$$
\begin{aligned}
V= & \frac{\lambda_{1}}{2}\left(\phi_{1}^{\dagger} \phi_{1}\right)^{2}+\frac{\lambda_{2}}{2}\left(\phi_{2}^{\dagger} \phi_{2}\right)^{2}+\lambda_{3}\left(\phi_{1}^{\dagger} \phi_{1}\right)\left(\phi_{2}^{\dagger} \phi_{2}\right)+\lambda_{4}\left(\phi_{1}^{\dagger} \phi_{2}\right)\left(\phi_{2}^{\dagger} \phi_{1}\right)+\frac{1}{2}\left[\lambda_{5}\left(\phi_{1}^{\dagger} \phi_{2}\right)^{2}+\text { h.c. }\right] \\
& -\frac{1}{2}\left\{m_{11}^{2}\left(\phi_{1}^{\dagger} \phi_{1}\right)+\left[m_{12}^{2}\left(\phi_{1}^{\dagger} \phi_{2}\right)+\text { h.c. }\right]+m_{22}^{2}\left(\phi_{2}^{\dagger} \phi_{2}\right)\right\} .
\end{aligned}
$$

$\delta_{1}=\operatorname{Arg}\left[\lambda_{5}^{*}\left(m_{12}^{2}\right)^{2}\right]$,
$\delta_{2}=\operatorname{Arg}\left[\lambda_{5}^{*}\left(m_{12}^{2}\right) v_{1} v_{2}^{*}\right]$
$\xrightarrow{\text { EWSB }}$

$$
\delta_{2} \approx \frac{1-\left|\frac{\lambda_{5} v_{1} v_{2}}{m_{12}^{2}}\right|}{1-2\left|\frac{\lambda_{5} v_{1} v_{2}}{m_{12}^{2}}\right|} \delta_{1}
$$

$h, H^{0}, A^{0} \rightarrow h_{1,2,3}$

| $\left(\begin{array}{cc}-s_{\alpha} c_{\alpha_{b}} & c_{\alpha} c_{\alpha_{b}} \\ s_{\alpha} s_{\alpha_{b}} s_{\alpha_{c}}-c_{\alpha} c_{\alpha_{c}}-s_{\alpha} c_{\alpha_{c}}-c_{\alpha} s_{\alpha_{b}} s_{\alpha_{c}} & c_{\alpha_{b} s_{\alpha_{c}}} \\ \hline s_{\alpha} s_{\alpha_{b}} c_{\alpha_{c}}+c_{\alpha} s_{\alpha_{c}} & s_{\alpha} s_{\alpha_{c}}-c_{\alpha} s_{\alpha_{b}} c_{\alpha_{c}} \\ c_{\alpha_{b}} c_{\alpha_{c}}\end{array}\right)$ |
| :---: | :---: | :---: |

CP mixing: $\alpha_{b} \& \alpha_{c}$ not independent


## Had \& Nuc Uncertainties

CPV \& 2HDM: Type II illustration

$$
\lambda_{6,7}=0 \text { for simplicity }
$$





## Present

$\sin \alpha_{b}: C P V$
scalar mixing

## Had \& Nuc Uncertainties

CPV \& 2HDM: Type II illustration

$$
\lambda_{6,7}=0 \text { for simplicity }
$$





## Present

## Challenge

$\sin \alpha_{b}: C P V$
scalar mixing

## Flavored EW Baryogenesis



Flavor basis (high T)

$$
\mathscr{L}_{\text {Yukawa }}^{\text {Lepton }}=-\overline{E_{L}^{i}}\left[\left(Y_{1}^{E}\right)_{i j} \Phi_{1}+\left(Y_{2}^{E}\right)_{i j} \Phi_{2}\right] e_{R}^{j}+\text { h.c. }
$$

Mass basis ( $T=0$ )
$\frac{m_{f}}{v} \kappa_{\tau}\left(\cos \phi_{\tau} \bar{\tau} \tau+\sin \phi_{\tau} \bar{\tau} i \gamma_{5} \tau\right) h$
Guo, Li, Liu, R-M, Shu 1609.09849
Chiang, Fuyuto, Senaha 1607.07316

## Flavored EW Baryogenesis



Jarlskog invariant

$$
J_{A}=\frac{1}{v^{2} \mu_{12}^{\mathrm{HB}}} \sum_{a, b, c=1}^{2} v_{a} v_{b}^{*} \mu_{b c} \operatorname{Tr}\left[Y_{c} Y_{a}^{\dagger}\right]
$$

T=0 Higgs couplings
$\operatorname{Im}\left(y_{\tau}\right) \sim \operatorname{Im}\left(J_{A}\right)$
EWBG CPV Source
$S C P V \sim \operatorname{Im}\left(J_{A}\right)$


Flavor basis (high T)
$\mathscr{L}_{\text {Yukawa }}^{\text {Lepton }}=-\overline{E_{L}^{i}}\left[\left(Y_{1}^{E}\right)_{i j} \Phi_{1}+\left(Y_{2}^{E}\right)_{i j} \Phi_{2}\right] e_{R}^{j}+$ h.c.
Mass basis ( $T=0$ )
$\frac{m_{f}}{v} \kappa_{\tau}\left(\cos \phi_{\tau} \bar{\tau} \tau+\operatorname{sPV} \rightarrow \tau \tau\right.$
Guo, Li, Liu, R-M, Shu 1609.09849
Chiang, Fuyuto, Senaha 1607.07316

## Flavored EW Baryogenesis




$$
\begin{aligned}
& \Delta \phi_{\tau} \sim 10^{\circ}: \\
& 3 a b^{-1} @ L H C 14
\end{aligned}
$$

