

Electric Dipole Moments: Phenomenology & Implications

M.J. Ramsey-Musolf

U Mass Amherst



AMHERST CENTER FOR FUNDAMENTAL INTERACTIONS

Physics at the interface: Energy, Intensity, and Cosmic frontiers

University of Massachusetts Amherst

<http://www.physics.umass.edu/acfi/>

ACFI Workshop, Amherst January 2015

Goals for this talk

- *Set the context for the workshop: What are key scientific questions & how do they fit in the broader context ?*
- *Illustrate the broader implications of present & prospective EDM searches*
- *Introduce terminology & notation*
- *Pose questions for hadronic structure theory*

Outline

- I. The BSM & NP context*
- II. Electric dipole moments*
- III. Questions for hadronic structure theory*
- IV. Outlook*

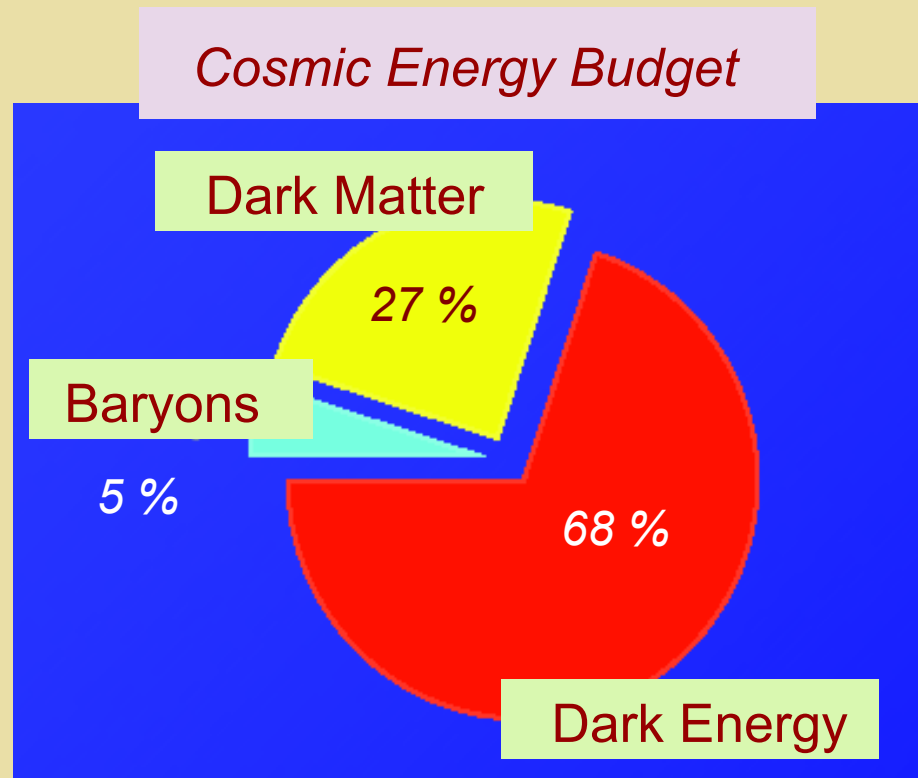
I. The BSM & NP Context

Scientific Questions

2007 NSAC LRP:

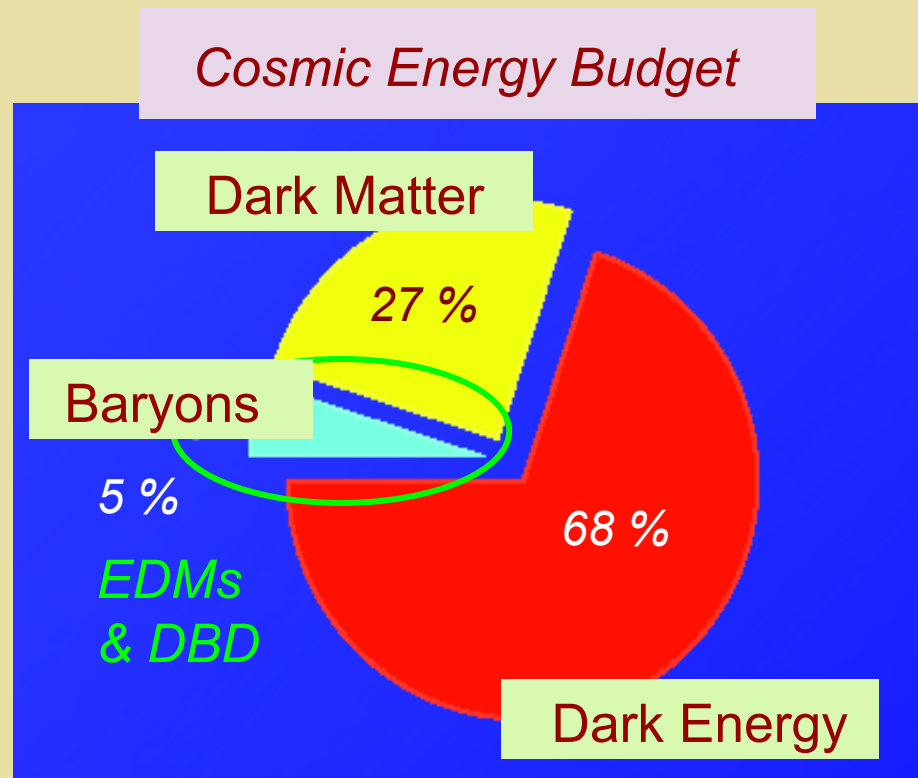
- *What are the masses of neutrinos and how have they shaped the evolution of the universe?*
- *Why is there more matter than antimatter in the present universe?*
- *What are the unseen forces that disappeared from view as the universe cooled?*

The Origin of Matter



Explaining the origin, identity, and relative fractions of the cosmic energy budget is one of the most compelling motivations for physics beyond the Standard Model

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Ingredients for Baryogenesis



- *B violation (sphalerons)*
- *C & CP violation*
- *Out-of-equilibrium or CPT violation*

Ingredients for Baryogenesis



Standard Model

BSM

- *B violation (sphalerons)*
- *C & CP violation*
- *Out-of-equilibrium or CPT violation*

✓

✓

✗

✓

✗

✓

Ingredients for Baryogenesis



Scenarios: *leptogenesis, EW baryogenesis, Affleck-Dine, asymmetric DM, cold baryogenesis, post-sphaleron baryogenesis...*

	<i>Standard Model</i>	<i>BSM</i>
• <i>B violation (sphalerons)</i>	✓	✓
• <i>C & CP violation</i>	✗	✓
• <i>Out-of-equilibrium or CPT violation</i>	✗	✓

Ingredients for Baryogenesis



Scenarios: leptogenesis,
EW baryogenesis, Affleck-
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baryogenesis, post-
sphaleron baryogenesis...

Testable

Standard Model

BSM

- B violation (sphalerons)
- C & CP violation
- Out-of-equilibrium or
 CPT violation

✓

✓

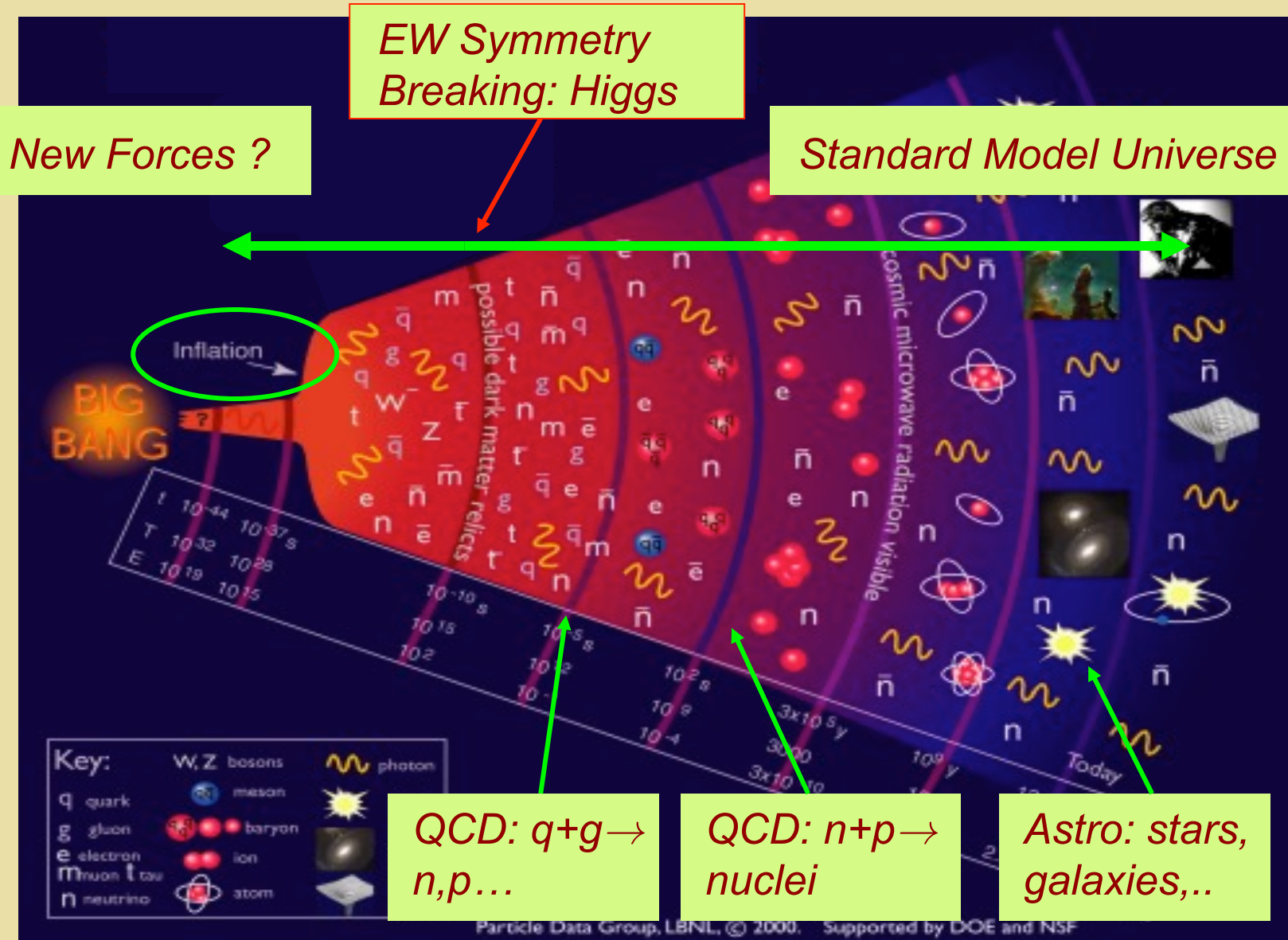
✗

✓

✗

✓

Symmetries & Cosmic History

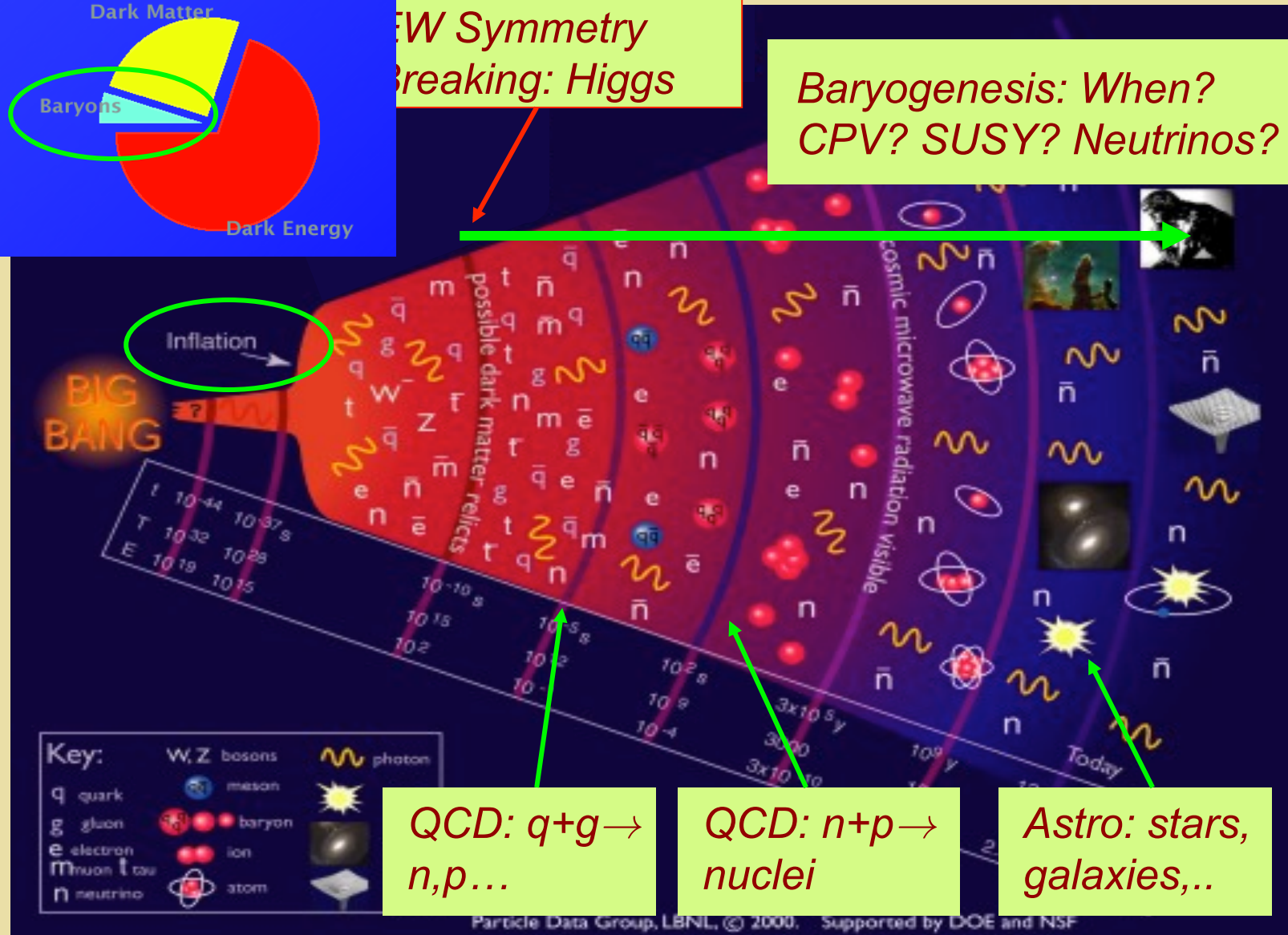


Symmetries & Cosmic History

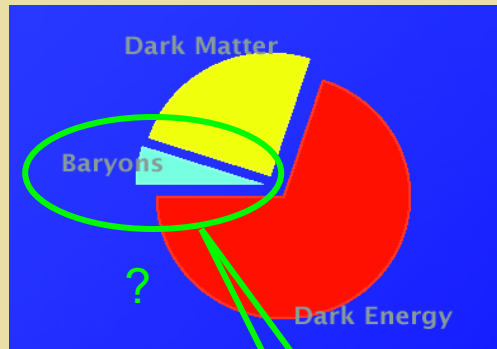


*EW Symmetry
Breaking: Higgs*

*Baryogenesis: When?
CPV? SUSY? Neutrinos?*

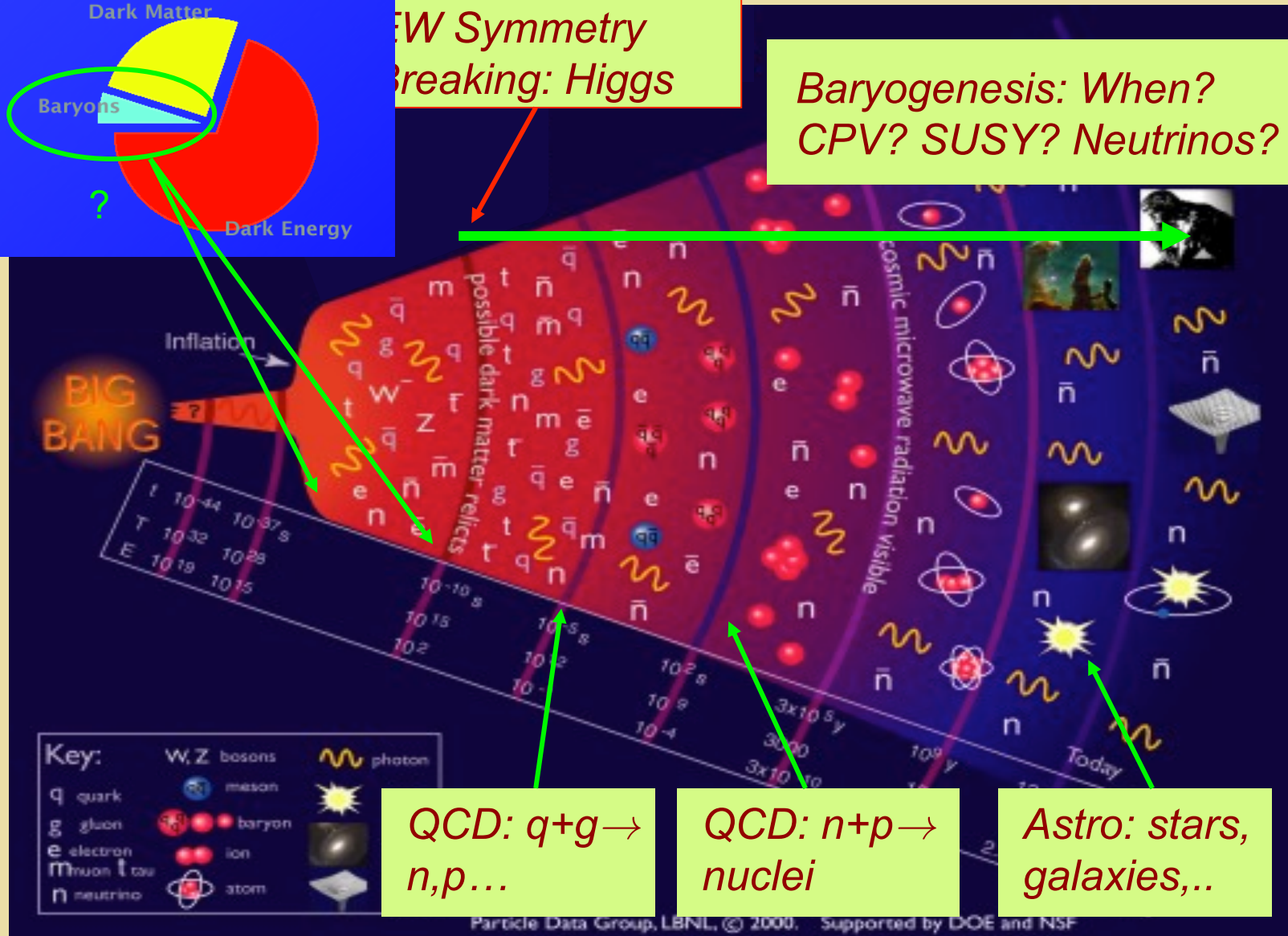


Symmetries & Cosmic History



*EW Symmetry
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Symmetries & Cosmic History



*EW Symmetry
Breaking: Higgs*

*Baryogenesis: When?
CPV? SUSY? Neutrinos?*

*Leptogenesis:
look for ingred's
w/ ν s: DBD, ν osc*

*EW Baryogenesis:
testable w/ EDMs +
colliders*

*QCD: $q+g \rightarrow$
 n, p, \dots*

*QCD: $n+p \rightarrow$
nuclei*

*Astro: stars,
galaxies,...*



Particle Data Group, LBNL, © 2000. Supported by DOE and NSF

Recent Results

- *Discovery of BEH-like scalar at the LHC*
- *Non-observation (so far) of sub-TeV particles at LHC*
- *New stringent limits on EDMs*

Recent Results

- *Discovery of BEH-like scalar at the LHC*
 - *Idea of ϕ -driven spontaneous EW symmetry breaking is likely correct*
- *Non-observation (so far) of sub-TeV particles at LHC*
 - *Sub-TeV BSM spectrum is compressed*
 - *Sub-TeV BSM is purely EW or Higgs portal*
 - *BSM physics lies at very different mass scale*
- *New stringent limits on EDMs*
 - *BSM CPV lies at high mass scale*
 - *BSM CPV doesn't talk directly to SM fermions*
 - *BSM CPV is flavor non-diagonal*

II. Electric Dipole Moments

- *Discovery potential & interpretation: need for searches in multiple systems*
- *Benchmark sensitivities: three examples*
- *Challenges & opportunities for hadronic & many-body theory*

EDMs: New CPV?

System	Limit (e cm)*	SM CKM CPV	BSM CPV
¹⁹⁹ Hg	3.1×10^{-29}	10^{-33}	10^{-29}
ThO	8.7×10^{-29} **	10^{-38}	10^{-28}
n	3.3×10^{-26}	10^{-31}	10^{-26}

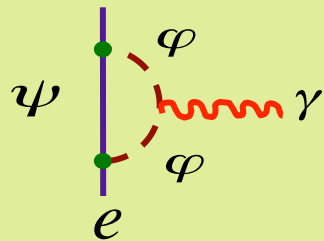
* 95% CL ** e⁻ equivalent

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Mass Scale Sensitivity



$$\sin\phi_{\text{CP}} \sim 1 \rightarrow M > 5000 \text{ GeV}$$

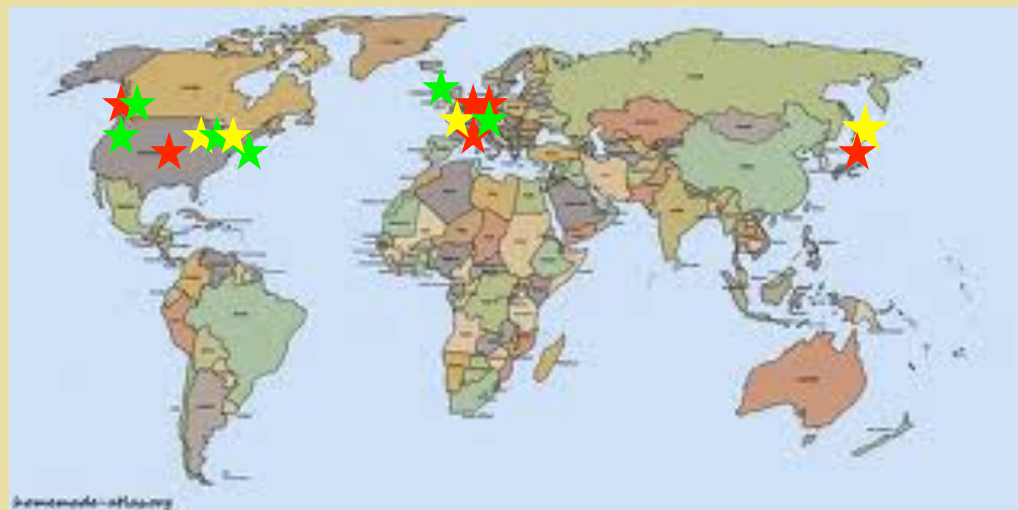
$$M < 500 \text{ GeV} \rightarrow \sin\phi_{\text{CP}} < 10^{-2}$$

EDMs: New CPV?

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Not shown:
muon



- ★ neutron
- ★ proton & nuclei
- ★ atoms

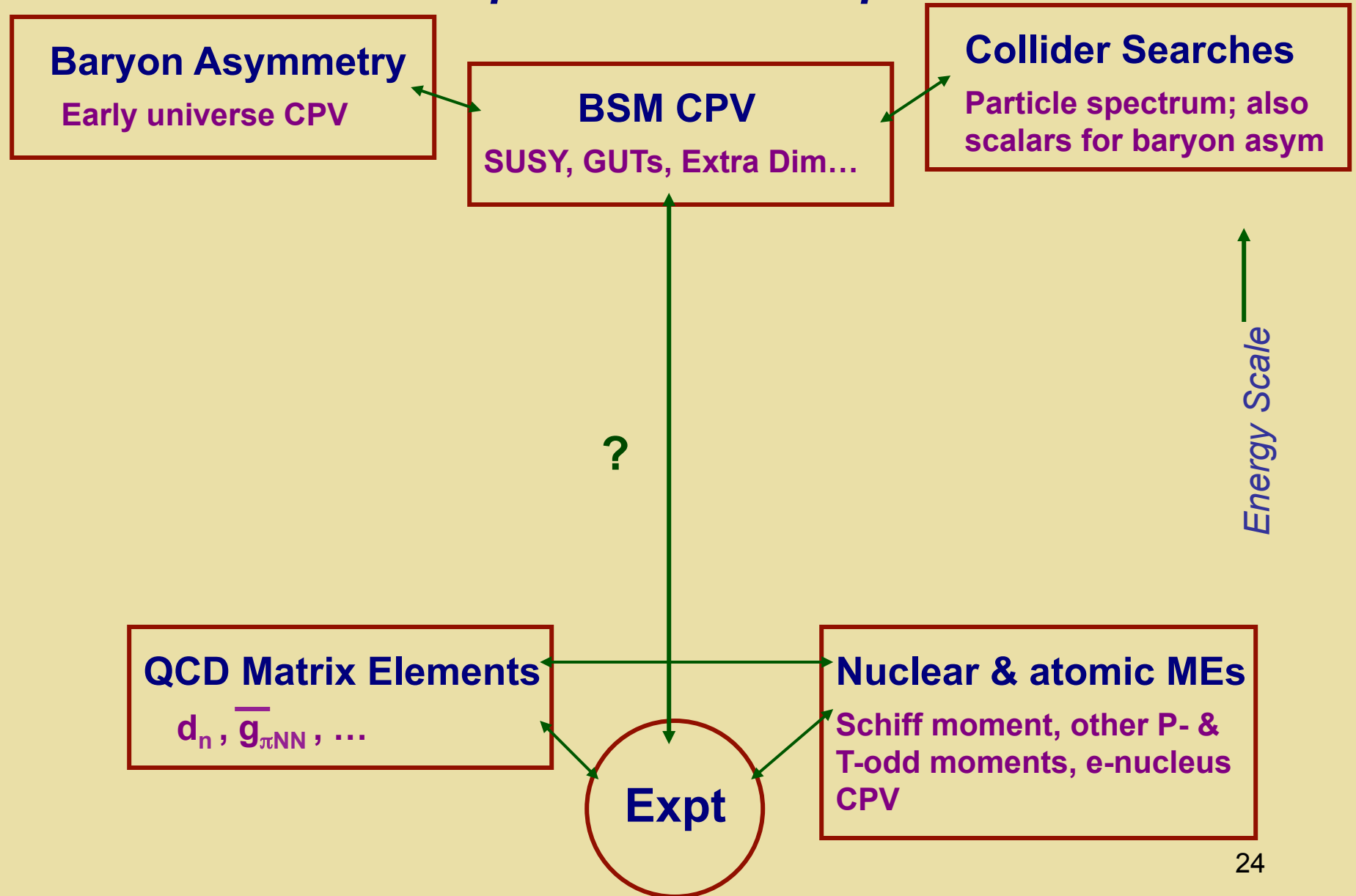
~ 100 x better sensitivity

Why Multiple Systems ?

Why Multiple Systems ?

Multiple sources & multiple scales

EDM Interpretation & Multiple Scales

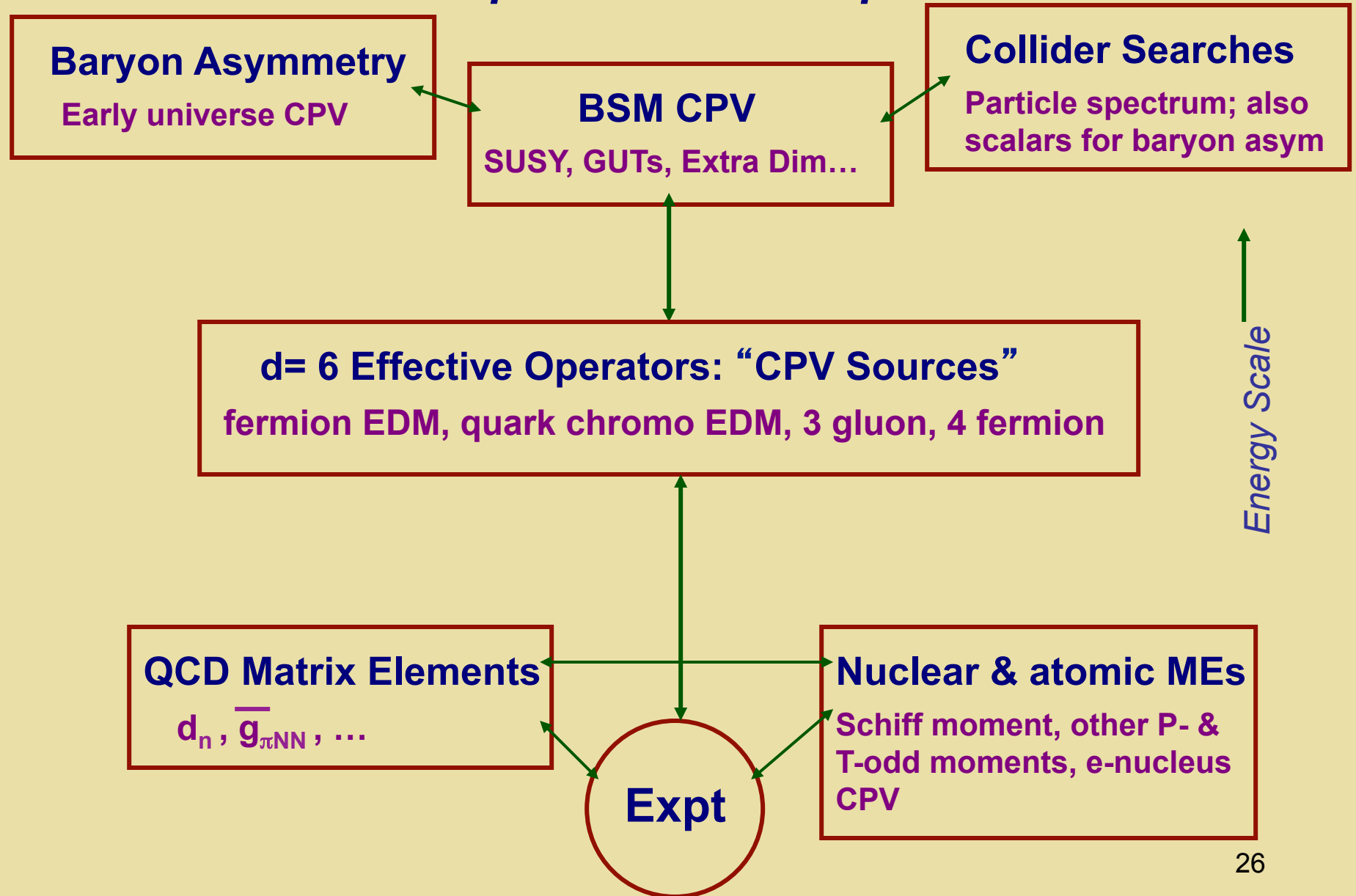


Effective Operators: The Bridge

$$\mathcal{L}_{\text{CPV}} = \mathcal{L}_{\text{CKM}} + \mathcal{L}_{\bar{\theta}} + \mathcal{L}_{\text{BSM}}^{\text{eff}}$$

$$\mathcal{L}_{\text{BSM}}^{\text{eff}} = \frac{1}{\Lambda^2} \sum_i \alpha_i^{(n)} O_i^{(6)} + \dots$$

EDM Interpretation & Multiple Scales



Wilson Coefficients: Summary

δ_f	<i>fermion EDM</i>	(3)
$\tilde{\delta}_q$	<i>quark CEDM</i>	(2)
$C_{\tilde{G}}$	<i>3 gluon</i>	(1)
C_{quqd}	<i>non-leptonic</i>	(2)
$C_{lequ, ledq}$	<i>semi-leptonic</i>	(3)
$C_{\varphi ud}$	<i>induced 4f</i>	(1)

12 total + $\overline{\theta}$

light flavors only (e,u,d)

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Complementary searches needed

BSM Origins

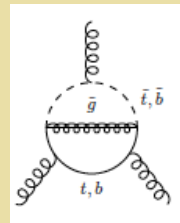
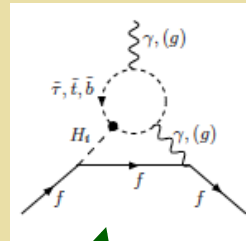
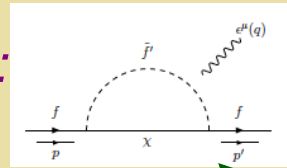
EDM: γff
 gff

CEDM:

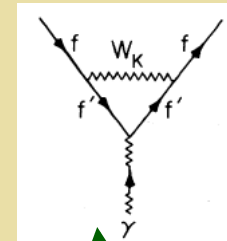
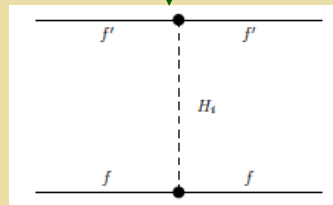
Weinberg ggg :

Four fermion

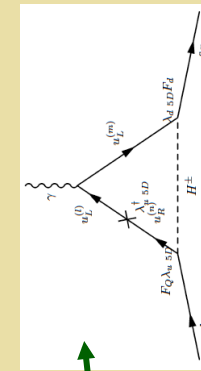
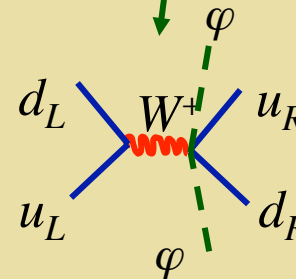
$udHH$



MSSM



LRSM



RS

Complementarity: Three Illustrations

- *CPV in an extended scalar sector (2HDM): “Higgs portal CPV”*
- *Weak scale baryogenesis (MSSM)*
- *Model-independent*

Higgs Portal CPV

Inoue, R-M, Zhang:
1403.4257

CPV & 2HDM: Type I & II

$\lambda_{6,7} = 0$ for simplicity

$$V = \frac{\lambda_1}{2}(\phi_1^\dagger \phi_1)^2 + \frac{\lambda_2}{2}(\phi_2^\dagger \phi_2)^2 + \lambda_3(\phi_1^\dagger \phi_1)(\phi_2^\dagger \phi_2) + \lambda_4(\phi_1^\dagger \phi_2)(\phi_2^\dagger \phi_1) + \frac{1}{2} \left[\lambda_5(\phi_1^\dagger \phi_2)^2 + \text{h.c.} \right] - \frac{1}{2} \left\{ m_{11}^2(\phi_1^\dagger \phi_1) + \left[m_{12}^2(\phi_1^\dagger \phi_2) + \text{h.c.} \right] + m_{22}^2(\phi_2^\dagger \phi_2) \right\}.$$

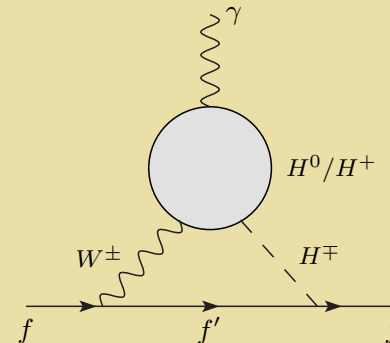
$$\begin{aligned} \delta_1 &= \text{Arg} \left[\lambda_5^* (m_{12}^2)^2 \right], \\ \delta_2 &= \text{Arg} \left[\lambda_5^* (m_{12}^2) v_1 v_2^* \right] \end{aligned}$$

EWSB

$$\delta_2 \approx \frac{1 - \left| \frac{\lambda_5 v_1 v_2}{m_{12}^2} \right|}{1 - 2 \left| \frac{\lambda_5 v_1 v_2}{m_{12}^2} \right|} \delta_1$$

Viable EWB & CPV:

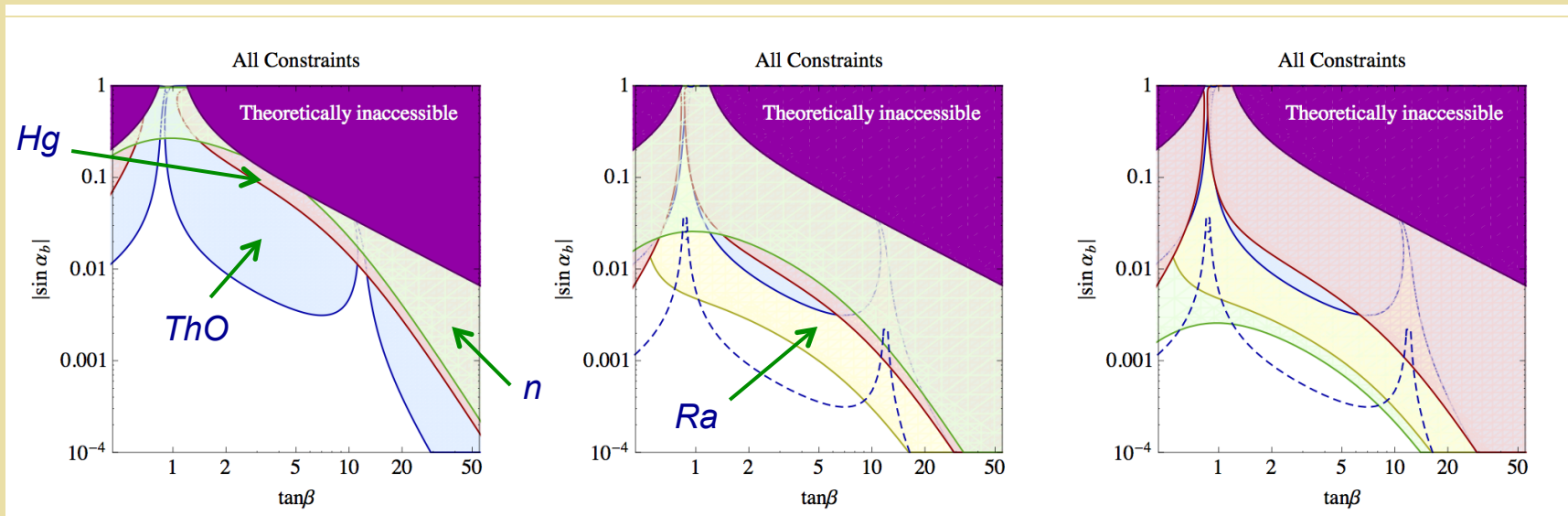
- EDMs are 2-loop
- CPV is flavor non-diag



Future Reach: Higgs Portal CPV

CPV & 2HDM: Type II illustration

$\lambda_{6,7} = 0$ for simplicity



Present

$\sin \alpha_b$: CPV
scalar mixing

Future:

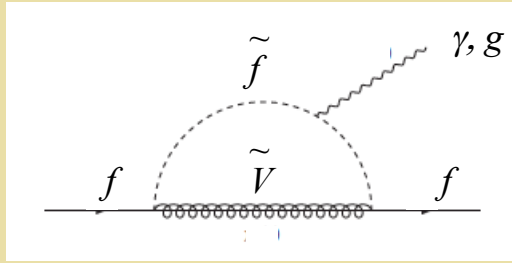
$d_n \times 0.1$
 $d_A(Hg) \times 0.1$
 $d_{ThO} \times 0.1$
 $d_A(Ra)$

Future:

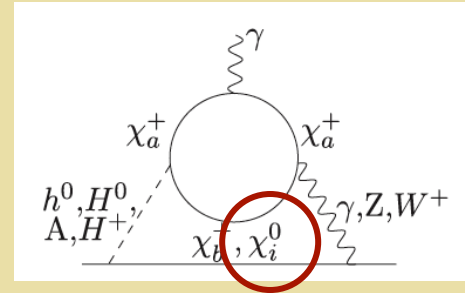
$d_n \times 0.01$
 $d_A(Hg) \times 0.1$
 $d_{ThO} \times 0.1$
 $d_A(Ra)$

Inoue, R-M, Zhang: 1403.4257

EDMs & EW Baryogenesis: MSSM

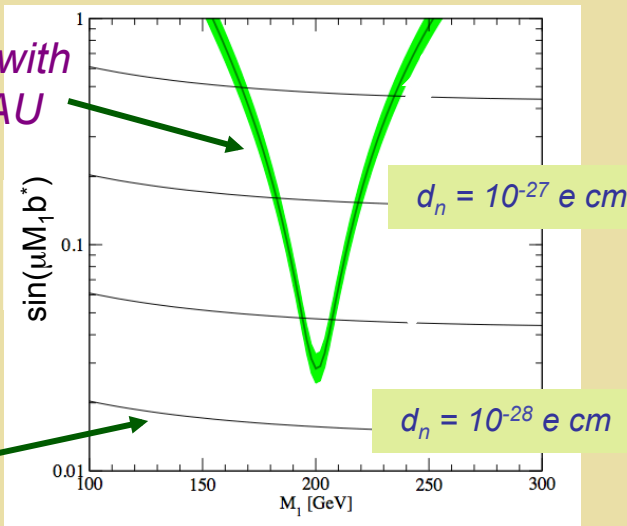


Heavy sfermions: LHC consistent & suppress 1-loop EDMs



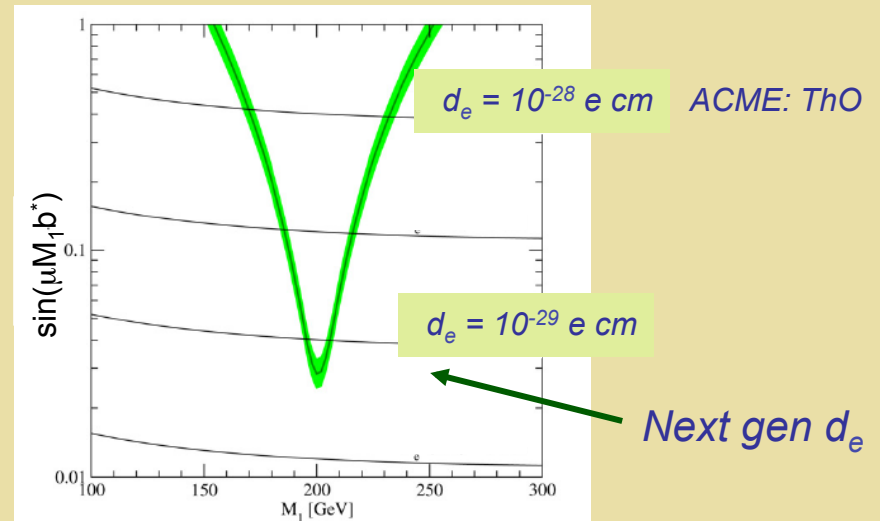
Sub-TeV EW-inos: LHC & EWB - viable but non-universal phases

Compatible with observed BAU



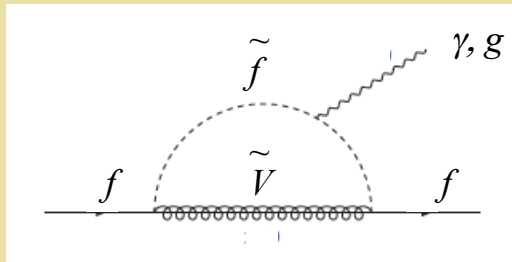
Next gen d_n

Li, Profumo, RM '09-'10

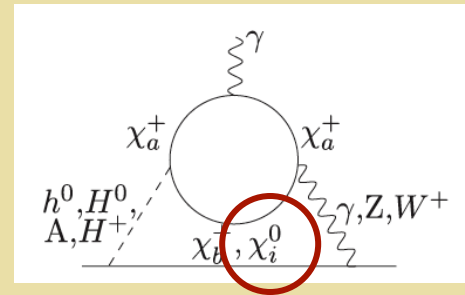


Next gen d_e

EDMs & EW Baryogenesis: MSSM

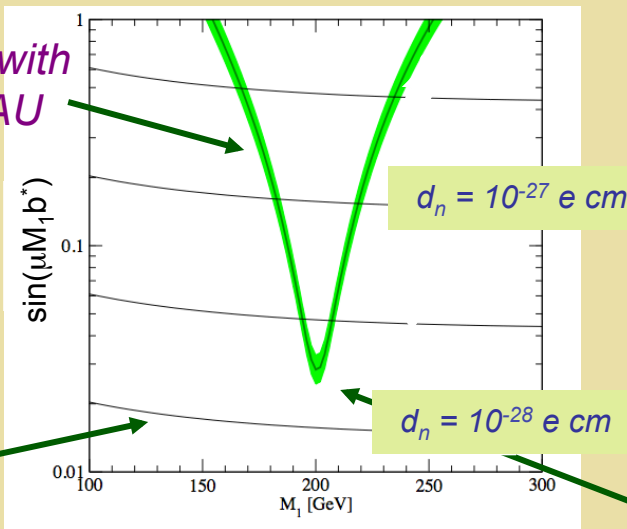


Heavy sfermions: LHC consistent & suppress 1-loop EDMs

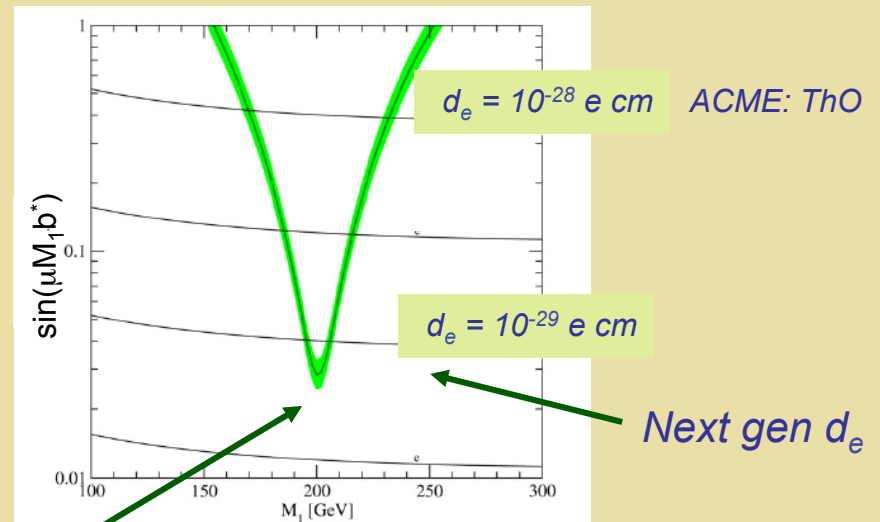


Sub-TeV EW-inos: LHC & EWB - viable but non-universal phases

Compatible with observed BAU



Li, Profumo, RM '09-'10



Compressed spectrum (stealthy SUSY)

Wilson Coefficients: Model Independent

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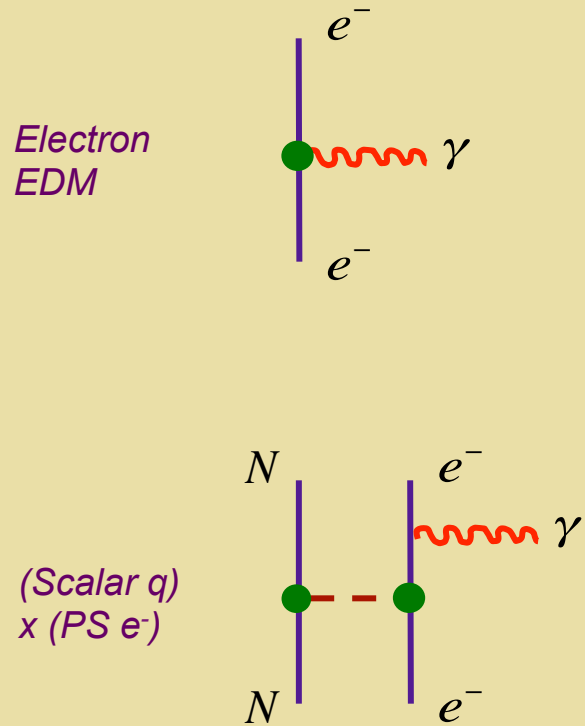
12 total + $\overline{\theta}$

light flavors only (e,u,d)

Global Analysis: Input

System	Year/ref	Result
Paramagnetic systems		
Cs	1989 [37]	$d_A = (-1.8 \pm 6.9) \times 10^{-24}$ e cm $d_e = (-1.5 \pm 5.6) \times 10^{-26}$ e cm
Tl	2002 [9]	$d_A = (-4.0 \pm 4.3) \times 10^{-25}$ e cm $d_e = (-6.9 \pm 7.4) \times 10^{-28}$ e cm
YbF	2011 [8]	$d_e = (-2.4 \pm 5.9) \times 10^{-28}$ e cm
ThO	2014 [7]	$\omega^{\mathcal{N}E} = 2.6 \pm 5.8$ mrad/s $d_e = (-2.1 \pm 4.5) \times 10^{-29}$ e cm $C_S = (-1.3 \pm 3.0) \times 10^{-9}$
Diamagnetic systems		
^{199}Hg	2009 [5]	$d_A = (0.49 \pm 1.5) \times 10^{-29}$ e cm
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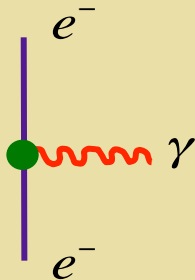
Paramagnetic Systems: Two Sources



Tl, YbF, ThO...

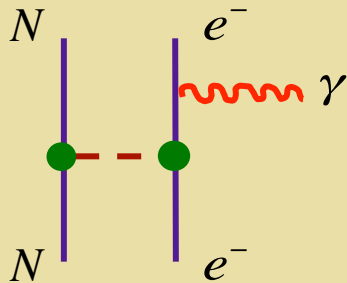
Paramagnetic Systems: Two Sources

Electron
EDM



$$d_f = -(1.13 \times 10^{-3} \text{ e fm}) \left(\frac{v}{\Lambda}\right)^2 Y_f \delta_f$$

(Scalar q)
 \times (PS e^-)

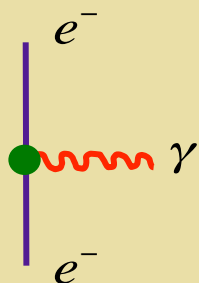


$$C_S^{(0)} = -g_S^{(0)} \left(\frac{v}{\Lambda}\right)^2 \text{Im } C_{eq}^{(-)}$$

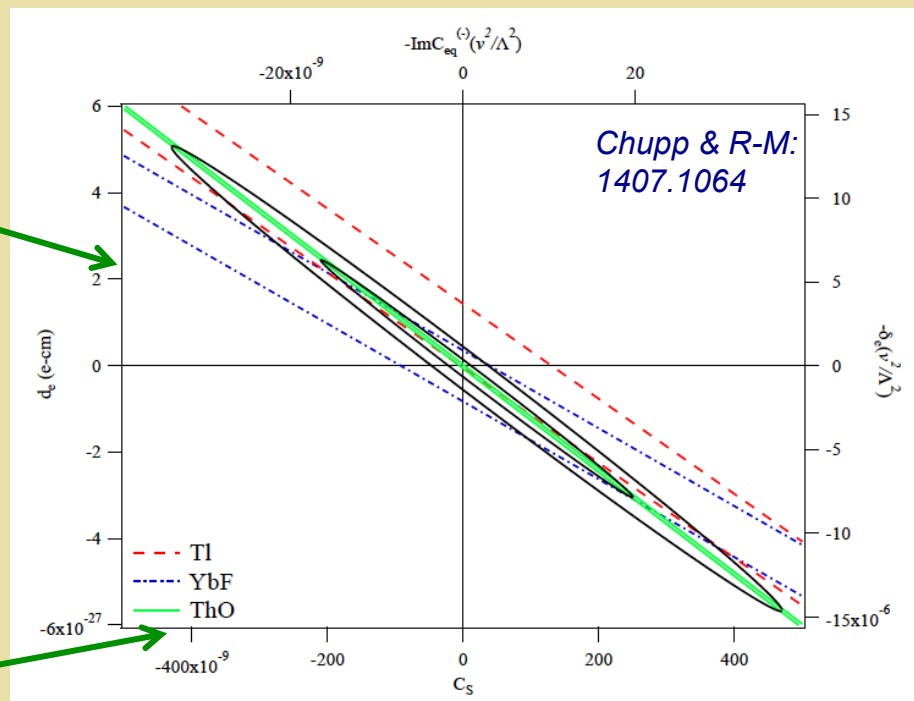
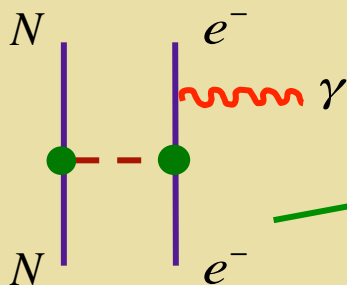
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Paramagnetic Systems: Two Sources

Electron
EDM

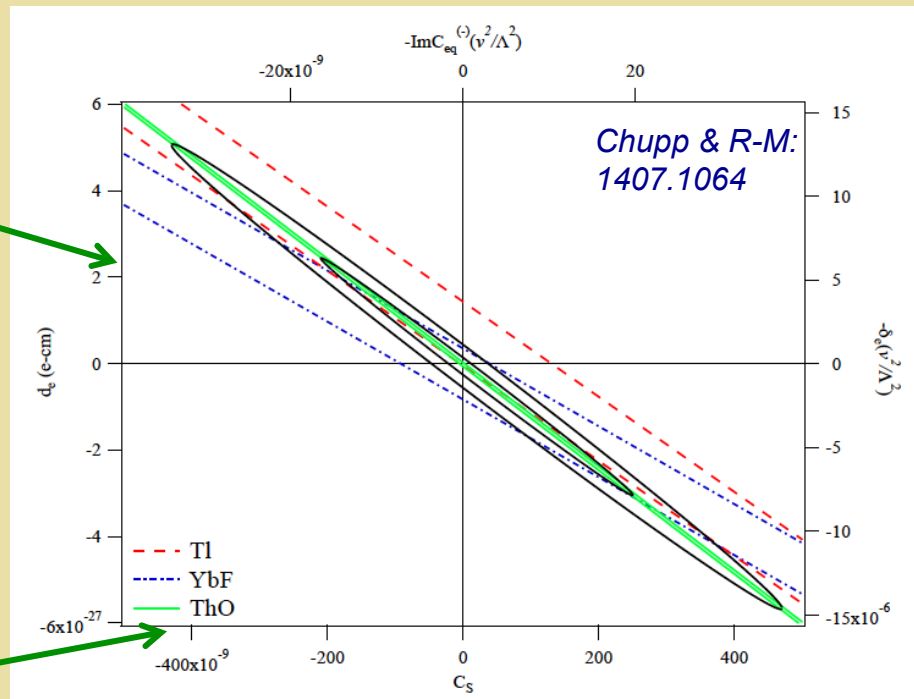
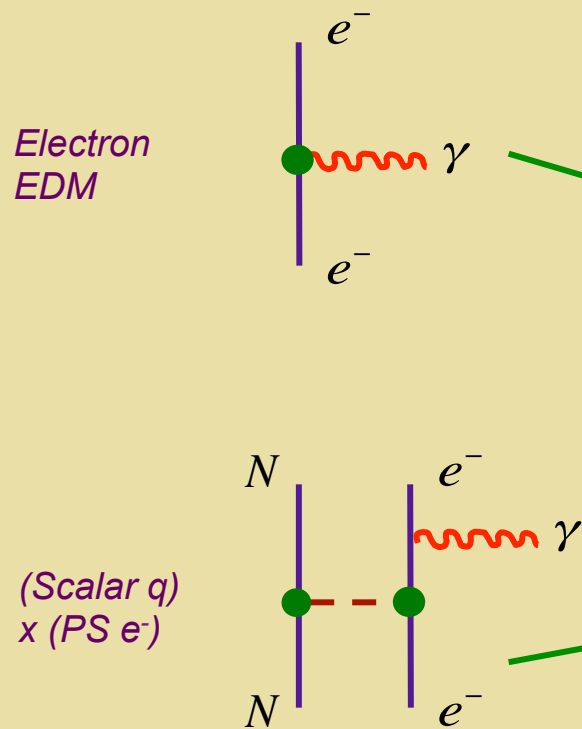


(Scalar q)
 \times (PS e^-)



Tl, YbF, ThO...

Paramagnetic Systems: Two Sources



$$\Lambda \gtrsim (1.5 \text{ TeV}) \times \sqrt{\sin \phi_{\text{CPV}}} \quad \text{Electron EDM (global)}$$

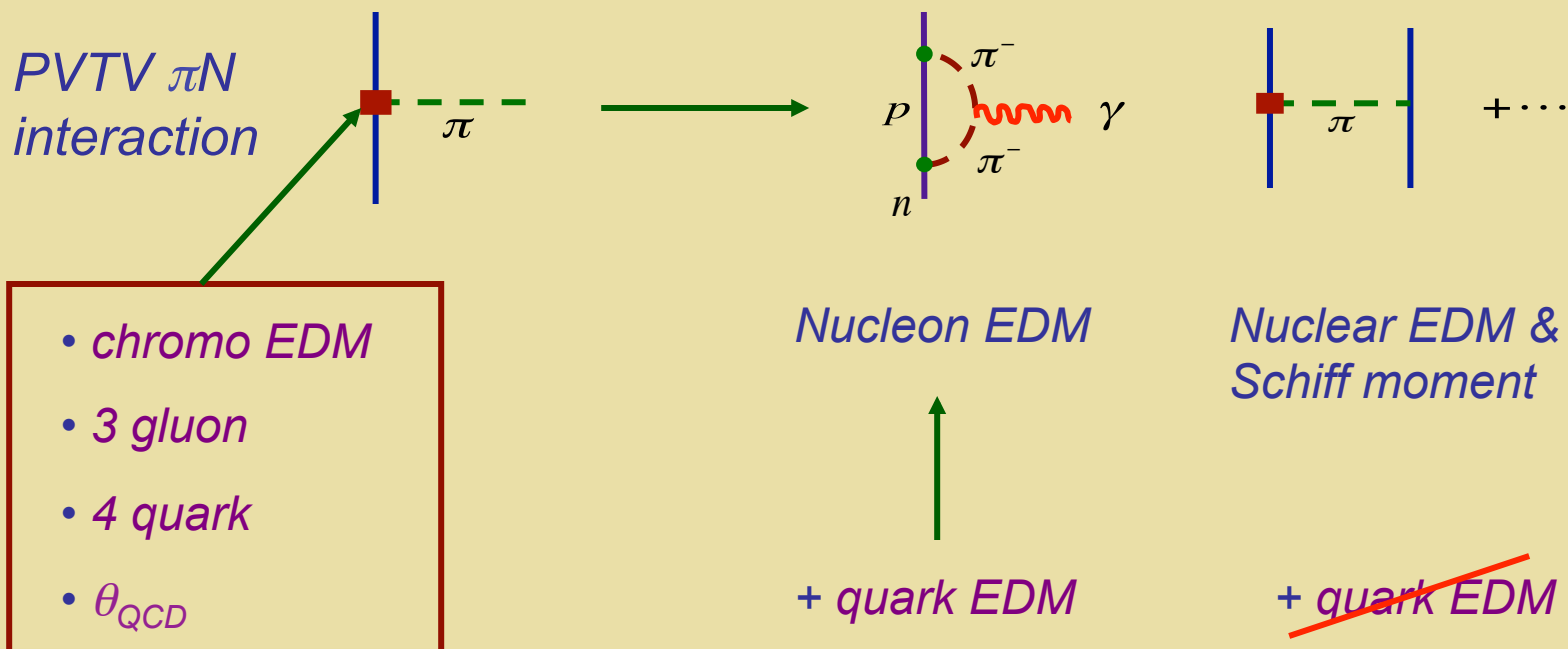
$$\Lambda \gtrsim (1300 \text{ TeV}) \times \sqrt{\sin \phi_{\text{CPV}}} \quad C_S \text{ (global)}$$

Tl, YbF, ThO...

Global Analysis: Diamagnetic Systems

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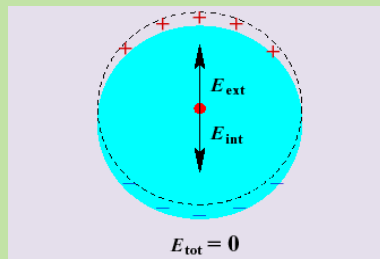
Hadronic CPV: Nucleons, Nuclei, Atoms



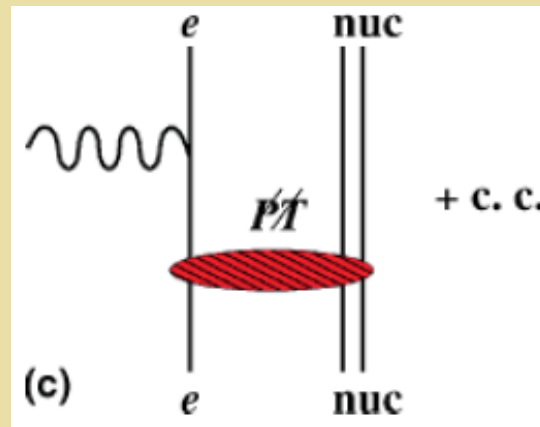
Neutron, proton & light nuclei (future), diamagnetic atoms

Diamagnetic Systems: P - & T -Odd Moments

Schiff Screening



Atomic effect from
nuclear finite size:
Schiff moment



Schiff moment, MQM, ...

EDMs of diamagnetic
atoms (^{199}Hg)

Diamagnetic Systems

Nuclear Moments

	PT	\cancel{PT}	$P\cancel{T}$	$\cancel{P}\cancel{T}$	
C_J	E	×	×	O	EDM, Schiff...
T^M_J	O	×	×	E	MQM....
T^E_J	×	O	E	×	Anapole...

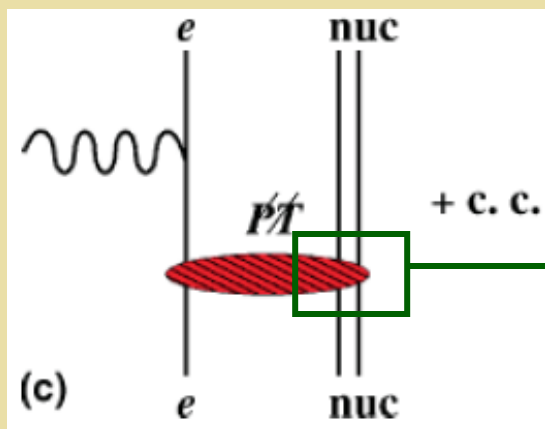
Diamagnetic Systems

Nuclear Moments

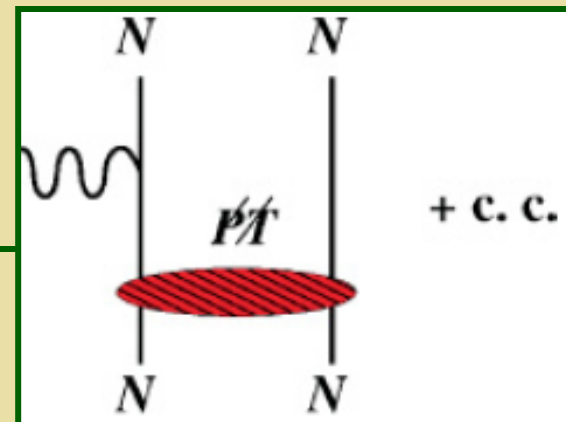
	PT	\cancel{PT}	$P\cancel{T}$	$\cancel{P}\cancel{T}$		
C_J	E	×	×	O	EDM, Schiff...	Nuclear Enhancements
T^M_J	O	×	×	E	MQM....	
T^E_J	×	O	E	×	Anapole...	

Nuclear Schiff Moment

Nuclear Enhancements



Schiff moment, MQM,...

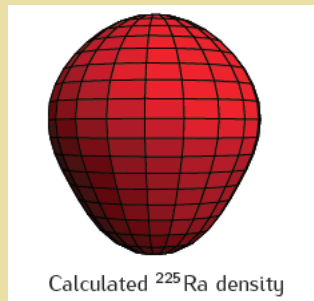


Nuclear polarization:
mixing of opposite parity
states by $H^{TVPV} \sim 1 / \Delta E$

EDMs of diamagnetic atoms (^{199}Hg)

Nuclear Schiff Moment

*Nuclear Enhancements:
Octupole Deformation*

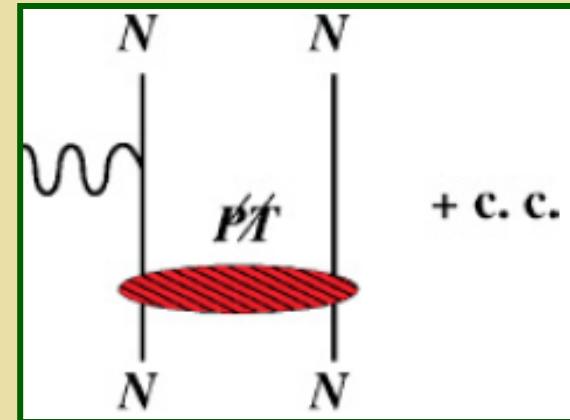


$$|\pm\rangle = \frac{1}{\sqrt{2}} (|\text{red}\rangle \pm |\text{blue}\rangle)$$

*Opposite parity states
mixed by H^{TVPV}*

“Nuclear amplifier”

EDMs of diamagnetic atoms (^{225}Ra)

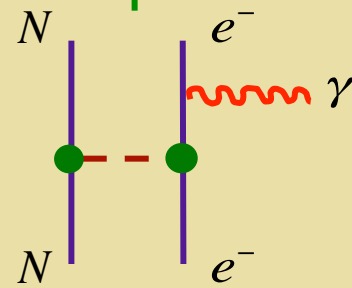


*Nuclear polarization:
mixing of opposite parity
states by $H^{\text{TVPV}} \sim 1 / \Delta E$*

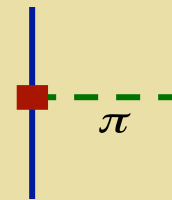
Thanks: J. Engel

Diamagnetic Global Fit

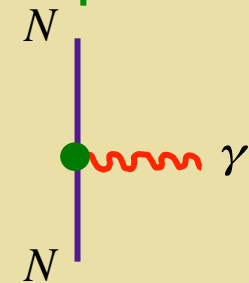
	$C_T \times 10^7$	$\bar{g}_\pi^{(0)}$	$\bar{g}_\pi^{(1)}$	\bar{d}_n (e-cm)
Exact solution	1.265	-6.687×10^{-10}	1.4308×10^{-10}	9.878×10^{-24}
Range from best values of α_{ij}	$(-7.6 - 9.5)$	$(-5.0 - 4.0) \times 10^{-9}$	$(-0.2 - 0.4) \times 10^{-9}$	$(-5.9 - 7.4) \times 10^{-23}$
Range from best values with $\alpha_{g_\pi^1}(\text{Hg}) = -4.9 \times 10^{-17}$	$(-7.6 - 8.4)$	$(-7.0 - 4.0) \times 10^{-9}$	$(0 - 0.2) \times 10^{-9}$	$(5.9 - 10.4) \times 10^{-23}$
Range from best values with $\alpha_{g_\pi^1}(\text{Hg}) = +1.6 \times 10^{-17}$	$(-9.2 - 12.4)$	$(-4.0 - 4.0) \times 10^{-9}$	$(-0.4 - 0.8) \times 10^{-9}$	$(-5.9 - 5.9) \times 10^{-23}$
Range from full variation of α_{ij}	$(-10.8 - 15.6)$	$(-10.0 - 8.1) \times 10^{-9}$	$(-0.6 - 1.2) \times 10^{-9}$	$(-12.0 - 14.8) \times 10^{-23}$



Tensor eq



TVPV πNN



Short distance d_n

Diamagnetic Global Fit

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Isoscalar
CEDM

$$\tilde{\delta}_q^{(+)} \left(\frac{v}{\Lambda} \right)^2 \lesssim 0.01 \quad \Lambda \gtrsim (2 \text{ TeV}) \times \sqrt{\sin \phi_{\text{CPV}}}$$

Caveat: Large hadronic uncertainty

Chupp & R-M:
1407.1064

Hadronic Matrix Elements

$$d_N = \alpha_N \bar{\theta} + \left(\frac{v}{\Lambda}\right)^2 \sum_k \beta_N^{(k)} \text{Im } C_k,$$

$$\bar{g}_\pi^{(i)} = \lambda_{(i)} \bar{\theta} + \left(\frac{v}{\Lambda}\right)^2 \sum_k \gamma_{(i)}^{(k)} \text{Im } C_k,$$

$$\bar{c}_i = \kappa_i \bar{\theta} + \left(\frac{v}{\Lambda}\right)^2 \sum_k \delta_i^{(k)} \text{Im } C_k,$$

$$\left(\frac{v}{\Lambda}\right)^2 \left[\beta_N^{qG} \text{Im } C_{qG} + \beta_N^{q\gamma} \text{Im } C_{q\gamma} \right] = e \tilde{\rho}_N^q \tilde{d}_q + \rho_N^q d_q = \left(\frac{v}{\Lambda}\right)^2 \left[e \tilde{\zeta}_N^q \tilde{\delta}_q + e \zeta_N^q \delta_q \right],$$

$$\left(\frac{v}{\Lambda}\right)^2 \left[\gamma_{(i)}^{qG} \text{Im } C_{qG} + \gamma_{(i)}^{q\gamma} \text{Im } C_{q\gamma} \right] = \tilde{\omega}_{(i)}^q \tilde{d}_q + \omega_{(i)}^q d_q = \left(\frac{v}{\Lambda}\right)^2 \left[\tilde{\eta}_{(i)}^q \tilde{\delta}_q + \eta_{(i)}^q \delta_q \right].$$

Hadronic Matrix Elements

Param	Coeff	Best value ^a	Range
$\bar{\theta}$	α_n	0.002	(0.0005–0.004)
	α_p	0.002	(0.0005–0.004)
$\text{Im } C_{qG}$	β_n^{uG}	4×10^{-4}	$(1 - 10) \times 10^{-4}$
	β_n^{dG}	8×10^{-4}	$(2 - 18) \times 10^{-4}$
\tilde{d}_q	$e\tilde{\rho}_n^u$	–0.35	–(0.09 – 0.9)
	$e\tilde{\rho}_n^d$	–0.7	–(0.2 – 1.8)
$\tilde{\delta}_q$	$e\tilde{\zeta}_n^u$	8.2×10^{-9}	$(2 - 20) \times 10^{-9}$
	$e\tilde{\zeta}_n^d$	16.3×10^{-9}	$(4 - 40) \times 10^{-9}$
$\text{Im } C_{q\gamma}$	$\beta_n^{u\gamma}$	0.4×10^{-3}	$(0.2 - 0.6) \times 10^{-3}$
	$\beta_n^{d\gamma}$	-1.6×10^{-3}	$-(0.8 - 2.4) \times 10^{-3}$
d_q	ρ_n^u	–0.35	(–0.17)–0.52
	ρ_n^d	1.4	0.7–2.1
δ_q	ζ_n^u	8.2×10^{-9}	$(4 - 12) \times 10^{-9}$
	ζ_n^d	-33×10^{-9}	$-(16 - 50) \times 10^{-9}$
$C_{\tilde{G}}$	$\beta_n^{\tilde{G}}$	2×10^{-7}	$(0.2 - 40) \times 10^{-7}$
$\text{Im } C_{\varphi ud}$	$\beta_n^{\varphi ud}$	3×10^{-8}	$(1 - 10) \times 10^{-8}$
$\text{Im } C_{quqd}^{(1,8)}$	β_n^{quqd}	40×10^{-7}	$(10 - 80) \times 10^{-7}$
$\text{Im } C_{eq}^{(-)}$	$g_S^{(0)}$	12.7	11–14.5
$\text{Im } C_{eq}^{(+)}$	$g_S^{(1)}$	0.9	0.6–1.2

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$\tilde{\delta}_q$ (CEDM)	$e\tilde{\zeta}_n^u$	8.2×10^{-9}	$(2 - 20) \times 10^{-9}$
	$e\tilde{\zeta}_n^d$	16.3×10^{-9}	$(4 - 40) \times 10^{-9}$
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Nuclear Matrix Elements

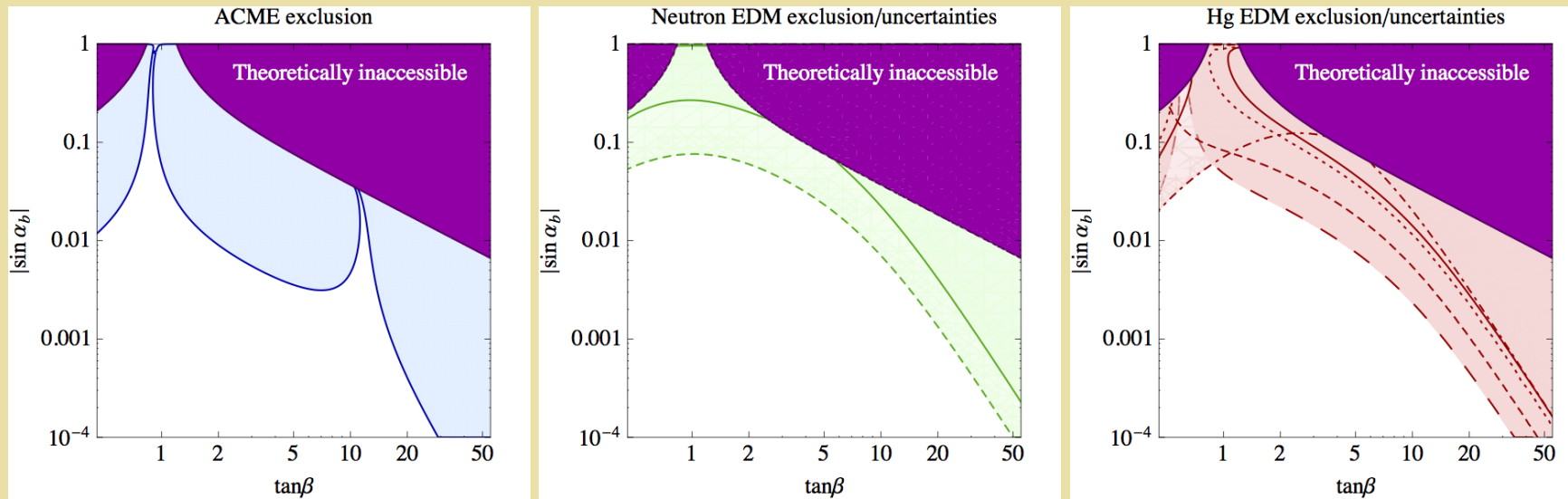
$$S = a_0 g \bar{g}_\pi^{(0)} + a_1 g \bar{g}_\pi^{(1)} + a_2 g \bar{g}_\pi^{(2)}$$

Nucl.	Best value		
	a_0	a_1	a_2
^{199}Hg	0.01	± 0.02	0.02
^{129}Xe	-0.008	-0.006	-0.009
^{225}Ra	-1.5	6.0	-4.0
Range			
	a_0	a_1	a_2
	0.005-0.05	-0.03-(+0.09)	0.01-0.06
	-0.005-(-0.05)	-0.003-(-0.05)	-0.005-(-0.1)
	-1-(-6)	4-24	-3-(-15)

Had & Nuc Uncertainties

CPV & 2HDM: Type II illustration

$\lambda_{6,7} = 0$ for simplicity



Present

$\sin\alpha_b$: CPV
scalar mixing

Inoue, R-M, Zhang: 1403.4257

III. Questions

- *What is the roadmap for reducing hadronic theory uncertainties for EDMs?*
- *What progress can be achieved through different approaches (lattice, DSE, EFT...)? Are they complementary ? If so, how ?*
- *What are the key conceptual and/or technical challenges that must be addressed make progress?*
- *Are there emerging new directions that call for further theoretical progress (e.g., few-body nuclear EDMs) ?*

IV. Outlook

- *Searches for permanent EDMs of atoms, molecules, hadrons and nuclei provide powerful probes of BSM physics at the TeV scale and above and constitute important tests of weak scale baryogenesis*
- *Studies on complementary systems is essential for first finding and then disentangling new CPV*
- *The interpretation of diamagnetic system EDMs (including the nucleon) is plagued by substantial hadronic and nuclear many-body uncertainties*
- *The advancing experimental sensitivity challenges hadronic structure theory to aim for an unprecedented level of reliability*